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Modulations of the Indian summer monsoon by the hot subtropical deserts: Insights

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from coupled sensitivity experiments

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24 **Abstract**

25 This study revisits the role of subtropical deserts in the Indian Summer Monsoon (ISM)
26 system by perturbing surface albedo over the subtropical deserts, to the west of the ISM domain
27 in different ways, using a state-of-the-art coupled model.

28 The analysis of up-to-date satellite datasets, atmospheric re-analyses and our control
29 coupled simulation suggests that the model broadly reproduces the radiation budgets close to re-
30 analyses and observed datasets. However, there are large uncertainties in the top-of-atmosphere
31 radiation budget over the Northern Hemisphere (NH) subtropical desert region during boreal
32 summer; while the model has a rather neutral radiation budget during boreal summer over the
33 Sahara Desert, the European Centre for Medium Range Weather Forecasts Interim reanalysis
34 show in contrast a radiative excess throughout the NH desert region and the up-to-date satellite
35 dataset has a clear negative radiation budget over north-eastern Sahara region and over Arabian
36 Peninsula.

37 Taking into account these uncertainties, our key finding is that by darkening the deserts
38 and arid regions to the west of ISM through a negative albedo perturbation in our coupled model,
39 the length and intensity of the rainy season over the Indian region are both significantly
40 increased with two well-defined rainfall anomaly maxima in May-June and September-October.
41 The ISM onset is advanced by one month and is characterized by a rapid northward propagation
42 of the rainfall band over the Indian domain.

43 Reversing the sign of our artificial albedo perturbation over the deserts in the model gives
44 an opposite response, highlighting the robust role of the subtropical deserts in the ISM system,
45 but the amplitude of the ISM response is also significantly larger, demonstrating nonlinearity in
46 the monsoon-desert relationship. Additional albedo perturbation experiments further demonstrate

47 that the whole hot subtropical deserts extending across Afro-Asian continents, and including the
48 Sahara, plays a key-role in the ISM response.

49 Finally, the modulations of the meridional tropospheric temperature gradient along with
50 stronger equatorial asymmetry of mean easterly shear and moisture distribution over the Indian
51 domain are key-factors for explaining the ISM response and its nonlinearity to the albedo
52 perturbations over the NH subtropical deserts. Further insights from moisture budget show that
53 the nonlinearity in advection moisture tendencies manifests in nonlinearity of the ISM response.

54

55 **Keywords:**

56 Indian summer monsoon; subtropical deserts; land albedo perturbations; radiation energy budget;
57 monsoon seasonality; abrupt monsoon transition

58

59 **1. Introduction**

60 Monsoon and desert coexist as twins at the northern hemisphere (NH) subtropical
61 latitudes of the African-Asian continent with their peak intensity during boreal summer (Rodwell
62 and Hoskins 1996; Wu et al. 2009). Two important examples are the Arabian-Iran-Thar desert
63 located just to the west of the Indian Summer Monsoon (ISM) system (Sikka 1997) and the
64 Sahara just to the north of the West African Monsoon (WAM) system (Lavaysse et al. 2009).
65 Consequently, relationships between monsoons and deserts have been widely investigated
66 (Ramage 1966; Charney 1975; Webster 1994; Rodwell and Hoskins 1996; Bollasina and Nigam
67 2011a,b; Biasutti et al. 2009; Lavaysse et al. 2009; Bollasina and Ming 2013; Vinoj et al. 2014;
68 Shekhar and Boos 2017; among many others).

69 The most conspicuous desert climate features include clear sky conditions, low rainfall,
70 high temperature, reduced soil moisture, lack of vegetation and high surface albedo year around
71 (Sikka 1997; Warner 2004). Radiative cooling of the atmospheric column prevails over these
72 regions and is compensated by upper level atmospheric subsidence. On the other hand, intense
73 solar heating of the surface during daytime and summer produces a strong SH flux into the lower
74 atmosphere, which creates near-surface low pressure systems usually referred as “heat” lows
75 (Ramage 1966; Warner 2004). The heat low over the Arabian-Iran-Thar desert is the deepest
76 surface low-pressure system in the Tropics during boreal summer. Similarly, the Saharan Heat
77 Low (SHL) over the Sahara is an important component of the WAM (Biasutti et al. 2009).
78 Furthermore, these two regional heat lows are embedded within a gigantic subtropical heat
79 trough across North Africa to West Asia during boreal summer. Here and after, we will refer
80 collectively to these arid regions as the “hot subtropical desert”, which is broadly defined as the
81 geographical region delimited by 20°W-75°E, 15°-40°N (as highlighted in Fig. 1a).

82 The traditional theory attributes the formation of subtropical deserts to the subsidence and
83 associated adiabatic heating in the descending branch of the Hadley circulation (Warner 2004).
84 This view is now recognized as imperfect (Yang et al. 1992; Rodwell and Hoskins 1996; Wang
85 2006; Wu et al. 2009) since the NH Hadley mean-circulation is the weakest during boreal
86 summer when South Asia receives abundant monsoon rainfall and aridity is the highest to its
87 west, over the hot subtropical desert, at about exactly the same latitudes.

88 Yang et al. (1992) and Webster (1994) have proposed the concept of a closed “Walker
89 type” circulation linking convection, upward motion and diabatic heating over South Asia to
90 subsidence over the hot subtropical desert in order to explain why these two climates coexist at
91 the same time and latitude over the African-Asian continent. However, this view has been
92 refuted by the seminal work of Rodwell and Hoskins (1996, 2001), which constitutes now the
93 accepted paradigm for explaining the coexistence of the hot subtropical desert and ISM (Wang
94 2006; Bollasina and Nigam 2011a; Tyrlis et al. 2013; Cherchi et al. 2014).

95 In this theory, the aridity over the hot subtropical desert is still associated with the
96 subsidence induced by ISM convective heating, but this subsidence is not explained by simple
97 zonal-mean Hadley-cell or meridional-mean Walker-cell arguments (Lindzen and Hou 1988;
98 Hoskins 1996; Rodwell and Hoskins 1996, 2001; Tyrlis et al. 2013; Cherchi et al. 2014). In
99 particular, Rodwell and Hoskins (1996), demonstrated that remote diabatic heating over South
100 Asia induces a Rossby wave pattern to its west and that the large-scale descent over the hot
101 subtropical desert and the eastern Mediterranean region during boreal summer results mainly
102 from the interaction between these westward propagating Rossby waves and the mid-latitude
103 westerlies. Their trajectory analysis further revealed that the subsiding air over the hot
104 subtropical desert is of mid-latitude origin. Subsequent studies showed additional contributions

105 through local longwave radiative cooling (which they called a “local diabatic enhancement
106 mechanism”) in strengthening the descent over the hot subtropical desert (Rodwell and Hoskins
107 2001). This “monsoon-desert mechanism” further highlights that the hot subtropical desert is a
108 rather “passive” recipient in the monsoon-desert relationship (Bollasina and Nigam 2011a).

109 From a different perspective, it is however generally agreed that the desert and arid
110 regions play an important role on the WAM and ISM systems (Charney 1975; Charney et al.
111 1977; Shukla and Mintz 1982; Sud and Fennessy 1982; Sud et al. 1988; Claussen 1997; Bonfils
112 et al. 2000; Douville et al. 2001; Xue et al. 2004; Yasunari et al. 2006; Xue et al. 2010; Bollasina
113 and Nigam 2011b; Bollasina and Ming 2013). The relationship between variabilities of the
114 WAM and SHL is for example well established (Haarsma et al. 2005; Biasutti et al. 2009;
115 Lavaysse et al. 2009). ISM is also known to be associated with atmospheric variability over the
116 arid regions to its west on a range of time scales (Ramage 1966; Smith 1986a,b; Mooley and
117 Paolino 1988; Parthasarathy et al. 1992; Krishnamurti et al. 2010; Saeed et al. 2011; Bollasina
118 and Nigam 2011b; Bollasina and Ming 2013; Vinoj et al. 2014; Rai et al. 2015). Bollasina and
119 Ming (2013) found that surface heating over the northwestern semi-arid areas alone determines a
120 realistic northwestward migration of the ISM at both the seasonal and intra-seasonal time scales
121 without any external forcing (e.g. seasonal variations of insolation) in their Atmospheric General
122 Circulation Model (AGCM). Vinoj et al. (2014), using also an AGCM, showed the short-term
123 (within a week) modulation of ISM rainfall through the dust-induced land surface heating over
124 North African and West Asian regions. Interestingly, these results coincided with a previous
125 result that the boreal summer heat low over West Asian desert regions may serve as an important
126 mechanism in controlling moisture transport into the ISM region (Smith 1986a,b; Mohalfi et al.
127 1998). More recently, Rai et al. (2015) and Chakraborty and Agrawal (2017) found a significant

128 correlation between ISM variability and pre-monsoon atmospheric variations over the arid
129 northwest India, Pakistan, Afghanistan and Iran regions. They argue that the heat low over these
130 regions drives the monsoon winds during the first part of boreal summer.

131 Collectively, these studies demonstrate that the changes of surface heating over the hot
132 subtropical desert can affect the ISM. Furthermore, they suggest that the heat lows are a key
133 element in the monsoon circulations. However, this would seem to partly contradict the idea that
134 the deserts and their associated heat lows are remotely forced by ISM through a Rossby wave
135 response (Rodwell and Hoskins 1996, Bollasina and Nigam 2011a).

136 Here, we seek to resolve this paradox/contradiction in a coupled modeling framework by
137 imposing different sets of surface land albedo over the hot subtropical desert to the west of the
138 South Asian domain in an ocean-atmosphere coupled model. These experiments aim at providing
139 new insights into the role of the arid regions in modulating the ISM rainfall and circulation.

140 Section 2 outlines the coupled ocean-atmosphere model, the design of our sensitivity
141 experiments and the validation datasets used in this study. Section 3 reviews the radiation
142 budget, both at the Top-Of-Atmosphere (TOA) and the surface, and the atmospheric circulation
143 during boreal summer with a special focus on the hot subtropical desert and ISM, as seen in
144 observations and our coupled model. Sections 4 and 5 describe the results of the sensitivity
145 experiments. Section 6 provides a discussion on the specific role of the Sahara in the ISM
146 response and addresses the robustness of our results with the help of another coupled model.
147 Section 7 summarizes our main conclusions.

148

149 **2. Coupled model, sensitivity experiments and validation datasets**

150 Unless otherwise specified, our analysis and diagnostics mostly focus on the boreal
151 spring-summer season from May to September.

152 2.a Climate coupled model

153 We used the SINTEX_F2 (SINTEX hereafter) coupled model for our control simulation
154 and subsequent coupled sensitivity experiments. Refer to Masson et al. (2012) for a complete
155 description of the version of the SINTEX model used here. The information regarding the
156 atmospheric, oceanic and land components can be summarized as follows. The atmosphere
157 model ECHAM5.3 is run at T106 spectral resolution, with 31 hybrid sigma-pressure levels
158 (Roeckner et al. 2003). The oceanic component is the NEMO (Madec 2008) at $0.5^\circ \times 0.5^\circ$
159 horizontal resolution, 31 vertical levels and with an interactive sea ice model. The atmosphere
160 and ocean exchange quantities such as heat and momentum fluxes every 2 hours, with no flux
161 adjustment or correction.

162 ECHAM5.3 uses a land surface scheme with prescribed white sky snow-free surface
163 albedo climatology, but without any seasonal cycle (Roeckner et al. 2003). Following Terray et
164 al. (2017), this old background albedo climatology has been updated by a seasonal varying snow-
165 free background albedo climatology over land estimated from the Moderate Resolution Imaging
166 Spectro-radiometer (MODIS) snow-free gap-filled white-sky (e.g. diffuse) albedo product
167 MCD43GF-v5 over the 2003-2013 period (Schaaf et al. 2011). For more technical details on the
168 original albedo parameterization in the ECHAM5.3 and on the updated version used here,
169 readers are referred to Roeckner et al. (2003) and Terray et al. (2017), respectively. The tropical
170 climate, including ISM, is better simulated by this updated SINTEX configuration (Terray et al.
171 2017).

172

173 2.b Coupled sensitivity experiments

174 Albedo, surface roughness and soil moisture play an important role in the simulation of
175 the hot subtropical desert and its relationship with the adjacent monsoon systems as illustrated by
176 numerous numerical studies focusing on the Holocene climate (Lofgren 1995; Bonfils et al.
177 2000) or the processes of desertification in the Sahel or South Asia regions (Charney 1975; Sud
178 and Fennessy 1982; Laval and Picon 1986; Claussen 1997; Sud et al. 1998; Snyder 2010;
179 Bollasina and Nigam 2011a,b; Pitman et al. 2011; Mahmood et al. 2014). These sensitivity
180 studies typically increase (decrease) the surface albedo (or the surface roughness length) directly
181 or indirectly (e.g. by changes in leaf area index and fractional green vegetation cover) in an
182 atmospheric model. The findings suggest that such surface modifications over the arid regions
183 alter the local climate and subsequently affect the adjacent monsoon systems through impacts on
184 atmospheric dynamics and rainfall (Pielke 2001; Bollasina and Nigam 2011b; Agrawal and
185 Chakraborty 2016). In a seminal paper, Charney (1975) suggested that desertification over the
186 Sahel is a consequence of a positive feedback produced by the increased albedo associated with
187 land degradation.

188 However, most of these previous modeling studies used stand-alone atmospheric or
189 regional models, which may not be sufficient to account for all the feedbacks affecting the
190 monsoon system. The climate sensitivity may be drastically modified because of the lack of
191 ocean-atmosphere coupling and the fact that Sea Surface Temperatures (SST) are often specified
192 as seasonally varying climatological values in these experiments. This is an important
193 shortcoming, especially for the monsoon regions (Wang et al. 2004, 2005, 2008; Wu and
194 Kirtman 2005; Kumar et al. 2005; Prodhomme et al. 2014, 2015).

195 Thus, our primary goal is to reinvestigate the remote impacts of changes in surface
196 albedo over the hot subtropical desert on the ISM with the help of our updated SINTEX
197 configuration. More precisely, we performed a total of 7 coupled experiments starting at the
198 same initial conditions, but differing only in the specifications of the background surface
199 (diffuse) albedo over the subtropical deserts (see Table 1).

200 First, a control simulation (referred throughout the manuscript as CTRL) of 110 years is
201 performed with the updated SINTEX configuration. We then conducted a series of six coupled
202 sensitivity experiments of 60 years each. In the first one (called here and after as
203 Desert_Arab_m20), the background land albedo has been artificially decreased by -20% over the
204 Arabian-Pakistan-Thar desert (15° - 40° N, 35° - 75° E as highlighted in Fig. 1d). A similar
205 experiment, but with an artificial increase of +20% over the same domain, has also been
206 performed (called here and after as Desert_Arab_p20). Third and fourth sensitivity experiments
207 are similar, but the artificial decrease/increase of background albedo concerns the whole hot
208 subtropical desert extending up to the Sahara in the west (called here and after as Desert_m20
209 and Desert_p20, respectively, See Fig. 1a and Table 1). Finally, in Section 6, we discussed the
210 results of sensitivity experiments with albedo perturbations only over the Sahara desert (15° -
211 40° N, 20° W- 35° E as highlighted in Fig. 1b. See Table 1 for further details.

212 Differences between the simulated climate in each experiment and CTRL were computed
213 (after removing the first ten years of the simulations), which we broadly refer to as “anomalous
214 responses” or simply “responses” to albedo perturbations over selected desert regions. Note that
215 using only the first 60 years in CTRL run or excluding 15 or 20 years from the various runs do
216 not change the results, consistent with our earlier investigations that 50 years of simulation is
217 sufficient to assess the response due to a given albedo perturbation (e.g. Terray et al. 2017).

218 A local statistical test is applied to the rainfall differences fields in order to assess the
219 statistical significance of the results. The statistical significance of the differences was estimated
220 through a permutation test with 9999 shuffles, and rainfall differences significant at the 95%
221 confidence level are indicated in the figures. More details about this statistical test are given in
222 Terray et al. (2003).

223

224 2.c Observed and reanalysis datasets

225 To validate model outputs, we make use of the following datasets: 1) monthly-
226 accumulated precipitation from the Global Precipitation Climatology Project (GPCP version 2.1;
227 Huffman et al. 2009) over the 1985-2014 period; 2) monthly-mean atmospheric variables (i.e.
228 surface skin temperature (TS), three dimensional circulation and mean sea-level pressure
229 (MSLP)) derived from the 2.5° ERA-Interim (ERAi) reanalysis produced by the European
230 Centre for Medium-Range Weather Forecasts ([http://apps.ecmwf.int/datasets/data/interim-full-](http://apps.ecmwf.int/datasets/data/interim-full-daily)
231 [daily](http://apps.ecmwf.int/datasets/data/interim-full-daily); Dee et al. 2011), over the 1979-2014 period and 3) the Clouds and the Earth's Radiant
232 Energy System Energy Balanced and Filled (CERES-EBAF edition 2.8; Kato et al. 2013) for
233 TOA and surface from the period 2001–2015 is also used to validate the radiation fluxes from
234 the model simulations and the ERAi reanalysis.

235

236 **3. The monsoon-desert system in observations, reanalysis data sets and coupled control** 237 **simulations**

238 We first provide a validation of SINTEX in this section. Since many studies have already
239 documented the mean and variability of ISM using SINTEX (e.g. Masson et al. 2012; Joseph et
240 al. 2012; Terray et al. 2012, 2016, 2017; Prodhomme et al. 2014, 2015, Crétat et al. 2016), we

241 focus on the Desert climate and its connections to ISM with the help of an energy budget
242 analysis and a tridimensional atmospheric view of the monsoon-desert system.

243 As a starting point, we begin with the radiation and energy budgets at TOA and the
244 surface over the desert region during boreal summer (Table 2). Note that the albedo values given
245 in Table 2 are defined as the ratio between reflected and incoming shortwave radiations at the
246 surface (or TOA), which differs from the prescribed background snow-free albedo in the model
247 (as discussed in Section 2.a). First, the results comprehensively highlight many of the desert
248 characteristics (Warner 2004), with the desert incurring large shortwave radiation loss (at both
249 TOA and surface) due to high surface albedo and large long-wave upward emission from the
250 surface due to the high surface temperature. Furthermore, the net long-wave radiation budget is
251 highly deficient both at TOA and surface, consistent with the facts that the atmosphere over the
252 desert is cloud-free with a reduced water vapor content (Neelin and Held 1987; Raymond 2000;
253 Wang and Dickinson 2013).

254 A conspicuous feature is that the desert has a net gain of radiative energy both at the
255 surface and TOA despite of the large upward long-wave emission during boreal summer, with
256 net radiation estimates of around 90 Wm^{-2} at the surface and between $16\text{-}30 \text{ Wm}^{-2}$ at TOA
257 depending on the datasets. Note that earlier studies usually reckon the deserts to be large-scale
258 radiative energy sinks on the annual mean time scale, with the Outgoing Long-wave Radiation
259 (OLR) exceeding the net solar radiation at TOA (e.g. Charney 1975; Wang 2006). Accordingly,
260 the TOA net radiative budget over the desert from ERAi turns out to be -8 Wm^{-2} , at the annual
261 time scale. The corresponding annual estimates from CERES-EBAF and CTRL also show a
262 radiative loss amounting to -19 and -10 Wm^{-2} , respectively. In contrast, during the boreal
263 summer, Table 2 indicates that the net radiation budget estimates at TOA over desert show a

264 radiative heating of about 30 Wm^{-2} in ERAi and CTRL and 16 Wm^{-2} in CERES-EBAF (see
265 Table 2). A look at the spatial maps (Fig. S1) however suggests large uncertainties in the
266 regional aspects of the TOA radiation budget over NH desert region during boreal summer. For
267 example, the couple model has a rather neutral radiation budget over the Sahara Desert (Fig.
268 S1c). But the reanalysis show a radiative excess throughout the NH desert region (Fig. S1b).
269 Note that earlier observational studies (using NIMBUS-7 radiation budget, Smith 1986a,b) also
270 showed surplus radiative energy over the Arabian desert which is again in contrast with the
271 observed estimates from CERES-EBAF (Fig. S1a). As per the other studies, the boundary layer
272 dust is observed to increase the short-wave absorption and this dust-induced heating overcomes
273 the long-wave radiative deficit over the Arabian desert, thus producing either a radiative neutral
274 or an energy surplus at TOA and strengthening of the heat low over there during boreal summer
275 (Ackerman and Cox 1982; Mohalfi et al. 1998).

276 Focusing now on the energy balance at the surface, Table 2 shows that approximately
277 75% of the net radiation gain for the surface is partitioned to sensible heat flux (SH), while about
278 20% contributes to latent heat flux (LH) in ERAi. SINTEX reasonably captures this surface
279 energy balance since the SH and LH represent, respectively, 85 and 10% of the available
280 radiative energy at the surface in CTRL. However, as expected, the LH flux estimates differ
281 significantly between ERAi and CTRL in absolute value, highlighting the large incertitude of
282 this component of the energy budget, even for the desert region where evaporation is small and
283 precipitation very sporadic (Wild et al. 2015).

284 It is also worth noting, finally, that the ERAi radiation estimates are more comparable to
285 CTRL than CERES-EBAF. Consistently, recent studies noted that ERAi contains significant
286 differences at the surface and TOA for radiative fluxes at the annual time scale, when compared

287 to CERES-EBAF (Wild et al. 2015; Rai et al. 2017). This highlights the uncertainties in state-of-
288 the-art atmospheric re-analyses. The radiation budget estimates specifically focusing on Afro-
289 Asian subtropical deserts are not so much, except a few studies (e.g. Haywood et al 2005;
290 Cherchi et al 2014; Blake et al 1983; Smith 1986a,b), which also showed the large errors in the
291 estimation of net radiation flux at TOA. These studies attributed the discrepancies to
292 uncertainties in the radiation parameterization code (e.g. absorption bands), which might also
293 result from errors in surface temperature, surface albedo and dust aerosol effects.

294 We performed similar budget diagnostics over ISM land domain (Table 3). Firstly, the
295 vegetated land surface over South Asia shows a much smaller surface albedo as compared to the
296 surface of the desert, but TOA albedo is stronger in all datasets. Consistently, net shortwave
297 radiation at the surface is much reduced by about 50 Wm^{-2} compared to the desert estimates,
298 highlighting the important role of clouds and a moist atmosphere. Secondly, the net long-wave
299 losses both at the surface and TOA over these cloudy regions (Table 3) are less compared to that
300 of arid regions, as the clouds and moisture tend to absorb the upward long-wave flux emitted
301 from the ground (Raymond 2000; Neelin and Held 1987). Finally as expected, the LH fluxes
302 largely dominates the SH fluxed over the ISM region.

303 We next looked at the seasonal evolution of monsoon rainfall over the Afro-Asian region
304 with a special focus on ISM (Fig. 1). In April, the Inter-Tropical Convergence Zone (ITCZ) is
305 located just south of the Equator over the Indian Ocean (IO) with two maxima, one off shore of
306 Sumatra and the second off the African coast, while the ITCZ already straddles the Equator over
307 the Atlantic Ocean (Fig. 1a). The commencement of rainy season in the IO occurs first over the
308 Sumatra region with the northward expansion of rainfall over Bay of Bengal and eastern Arabian
309 Sea in May (Fig. 1b). On the other hand, during this month, the northward migration of the ITCZ

310 over the western IO and the African sector is very slow (e.g. about 1° of latitude). Afterwards,
311 the monsoon convection spreads northwards leading to the ISM onset over the regions
312 neighboring southern India during June (Fig. 1c). The monsoon penetrates further to central
313 Indian landmass in the subsequent months while the ITCZ is located more southward and looks
314 almost stationary in the African region (Fig. 1d-e).

315 CTRL shows a very similar rainfall evolution. The spatial correlations between CTRL
316 and GPCP are always above 0.73 for the tropical domain (i.e. within 20°S to 35°N ; see Fig. 1)
317 from April to September. However, the simulated precipitation exhibits some well-known biases
318 with a more southward position of the ITCZ in the African sector, excessive rainfall over the IO
319 north of the equator and deficient rainfall over the eastern equatorial IO and the central Indian
320 landmass (Joseph et al. 2012; Terray et al. 2012; Prodhomme et al. 2014).

321 To set the stage for further understanding on the monsoon-desert system, Figure 2 shows
322 the vertical structure of vertical velocity and horizontal divergence, along a pressure-longitude
323 plane averaged over 15° - 30°N from May to September in ERAi and CTRL. The southern limit
324 of the domain is set at 15°N since this is approximately the boundary between the rainy and
325 desert regions in the Afro-Asian sector during boreal summer (Fig. 1).

326 As expected, ERAi shows centers of strong summertime ascent (see shading in Fig. 2)
327 with a marked seasonal evolution in the east, contrasting with well-defined subsidence from 200
328 to about 600-hPa to its west, especially during May-June (see Fig. 2a-b). The strong ascent,
329 extending throughout the troposphere, is situated over the Indian landmass and Bay of Bengal
330 (Indo-BoB region, 70° - 90°E) and reaches a maximum peak at July (Fig. 2a-c), thus, consistently
331 following the seasonal evolution of ISM rainfall (Fig. 1). As a thermodynamic response to the
332 vertical motion induced by diabatic heating (Neelin and Held 1987), there is low-level

333 convergence (see contours in Fig. 2) with associated low-level westerlies (not shown), and
334 upper-level divergence associated with upper-level easterlies (not shown) over the ISM domain,
335 thus resembling a baroclinic structure during the whole boreal summer (Jiang et al. 2004).

336 In May, one can also detect several well defined regional heat lows showing convergence
337 and ascending motion limited to lower levels (e.g. between the surface and 600-hPa) consistent
338 with the dry convection affecting these arid regions. This near-surface upward motion is also
339 quickly damped due to the counteracting effect of the prominent mid-to-upper level subsidence
340 between TOA and about 700-hPa (Fig. 2a). However, two interesting points are (i) the notable
341 abatement of this upper-level subsidence after May (Fig. 2b-c), coinciding with the northward
342 shift of the monsoon activity over the Indo-BoB longitudes (Fig. 1c-e), and (ii) the persistence of
343 the heat lows in July-August despite the surface temperature has significantly decreased at this
344 time (figure not shown). Furthermore, there is a tendency for the regional heat lows to merge
345 together from May to July leading to the formation of a gigantic heat trough after the sudden
346 attenuation in the upper level subsidence (Fig. 2a-c). CTRL shows a similar evolution of the
347 vertical velocity and convergence fields (Fig. 2f-j). The above description highlights the forcing
348 of the hot subtropical desert by the South Asian monsoon (Rodwell and Hoskins 1996, 2001;
349 Bollasina and Nigam 2011a), but, surprisingly, also a possible relaxation of this forcing after the
350 ISM onset in both ERAi and CTRL.

351 We extended our analysis in Figure 2, by examining the seasonal evolution of vertical
352 motion and horizontal divergence over 30° - 40° N (Fig. 3). Upward motion can still be noticed
353 eastward of 80° E over these latitudes, which correspond to the high Tibetan plateau, during the
354 whole ISM period. Distinct and strong descent centers can also be distinguished: one over west
355 Asia (in between 60° - 70° E) and another one over the eastern Mediterranean region (in between

356 10°-40°E), with the later one fully consistent with the monsoon-desert paradigm (Rodwell and
357 Hoskins 1996, 2001). In fact, the entire troposphere over the eastern Mediterranean sector is
358 under a strong descent regime. Interestingly, the subsidence over this region intensifies from
359 May to July (Fig. 3a-c and f-h), which is in contrast with what happened over the south part of
360 the hot subtropical desert where subsidence decreases over the same time interval (Fig. 2a-c and
361 f-h). The subsidence weakens only at a slow pace in August, with an accelerated weakening
362 thereafter. This evolution over the northern part of the desert domain (e.g. 30°-40°N) is again
363 well reproduced by CTRL (Fig. 3f-j).

364 In summary, the aforementioned features highlight the very different vertical structure of
365 the atmospheric circulation over the southern and northern regions of the hot subtropical desert.
366 They also qualify the coupled model for a reasonable representation of the monsoon-desert
367 system over the Afro-Asian region, especially the vertical structure between northern and
368 southern boundaries of the hot subtropical desert, the heat lows over the subtropical deserts as
369 well as the vertical atmospheric structure associated with the ISM, which are the features
370 believed to be important for further understanding of the relationships between ISM and the
371 deserts in the next sections.

372

373 **4. ISM response to reduced surface albedo over the hot subtropical desert**

374

375 In this section, we focus on the responses resulting from an artificial decrease in surface
376 albedo over the whole hot subtropical desert and the Arabia-Middle-East desert in the coupled
377 model (e.g. the Desert_m20 and Desert_Arab_m20 experiments in Table 1).

378

379 4.1 Energy budget and ISM rainfall response

380 Table 4 shows the energy budgets at TOA and the surface over the hot subtropical desert
381 in Desert_m20 and over the Arabia and Middle-East desert in Desert_Arab_m20, along with
382 their differences from CTRL. First, the results comprehensively highlight the changes in
383 radiative forcing caused by a surface albedo perturbation over the deserts. The surface albedo
384 over the deserts decreases by -19 % in the two experiments, e.g. by about the same amount as
385 imposed in the background land albedo in these experiments (see Table 1). As expected, the
386 amount of shortwave radiation absorbed by the surface increases (by about $+31 \text{ Wm}^{-2}$), but also
387 at TOA (about $+25 \text{ Wm}^{-2}$). This results in surface warming (of about 3K) and stronger surface
388 upward long-wave emission (about -19 Wm^{-2}). However, this radiative loss subsequently gets
389 offset by an even larger increase of downward long-wave flux, as indicated by the positive value
390 of net long-wave radiation budget at the surface in Desert_m20 (e.g. $+6 \text{ Wm}^{-2}$). This
391 demonstrates the existence of a positive feedback amplifying the original surface warming due
392 solely from the surface albedo perturbation. The LH and SH latent fluxes also increase, with the
393 later one dominating over the other (-8 and -29 Wm^{-2} , respectively). However, both the LH and
394 rainfall responses dominate the SH response in relative terms (the relative increases in LH and
395 rainfall are 89% and 160%, respectively, while it is only 39% in SH flux), suggesting a
396 significant increase of the water vapor content of the atmosphere over the hot subtropical desert
397 in the experiments with a possible effect on the net long-wave radiation budget at the surface as
398 described above.

399 The net radiation budget at TOA is even more positive than at surface in Desert_m20
400 (about $+39 \text{ Wm}^{-2}$ at TOA compared to $+37 \text{ Wm}^{-2}$ at the surface) leading to a modest radiative
401 warming of the whole atmospheric column over the hot subtropical desert (about $+2 \text{ Wm}^{-2}$). The
402 Desert_Arab_m20 experiment leads to similar results over the Arabia and Middle-East desert

403 (Table 4). The basic inference here is that by perturbing the surface albedo over the subtropical
404 deserts, we not only change their surface thermal characteristics, but also somewhat modify the
405 whole tropospheric column above the hot subtropical deserts as we will further demonstrate
406 below.

407 This excess of energy over the deserts leads to modifications of the energy budget over
408 the ISM region and an increase of ISM rainfall in Desert_m20 or Desert_Arab_m20 (Table 5).
409 Albedo over the ISM landmass remains unchanged at the surface, but increases at TOA in both
410 experiments (3% and 2%, respectively). This suggests increased cloud cover and abundant
411 moisture consistent with the enhanced rainfall in the two experiments (+2 and +1 mm/day in
412 Desert_m20 and Desert_Arab_m20, respectively; corresponding to an increase of 34% and 16%,
413 respectively). Accordingly, there is a modest decrease in net solar radiation both at TOA and
414 surface over the ISM region (e.g. -13 Wm^{-2} at TOA and -16 Wm^{-2} at surface in Desert_m20 for
415 example; see Table 5) in agreement with a cooling of the land surface (e.g. -0.7°C and -0.3°C in
416 Desert_m20 and Desert_Arab_m20, respectively; see Table 5). This shortwave radiation loss, is
417 only partly compensated by an increase of the net long-wave radiation (e.g. $+9 \text{ Wm}^{-2}$ and $+4$
418 Wm^{-2} , respectively at TOA and surface in Desert_m20; see Table 5), again in agreement with a
419 more cloudy and moist atmosphere over the Indian landmass in both Desert_m20 and
420 Desert_Arab_m20. Finally, the enhanced ISM rainfall is associated with an increase of LH flux,
421 but again the changes are more significant in Desert_m20 than Desert_Arab_m20.

422 Figures 4 (on the time-latitude evolution of monthly precipitation) and S2 (on the spatial
423 evolution of rainfall response) further illustrate the significant differences in ISM responses
424 between the two experiments. A notable feature in both the experiments is an early ISM onset
425 with a spatially well-organized rainfall band between 5° - 15°N occurring as soon as in May (see

426 Fig. 4b-c and Fig. S2b). The ITCZ, usually located over the equator during May, has already
427 moved to the north, thus occupying the eastern Arabian Sea and then extending to BoB region in
428 both the experiments (compare Figs. 1b and S2b). However, the Desert_m20 simulates a much
429 faster northward rainfall migration and more intense precipitation anomalies, with two well-
430 defined maxima in May-June and September-October, leading to a significant increase of the
431 ISM length (Fig. 4b). In contrast, the rainfall anomalies in Desert_Arab_m20 quickly fade away
432 after May-June (Fig. 4c). This suggests that some positive feedbacks are at work when we
433 considered the whole hot subtropical desert in the experiments.

434

435 4.2 Tropospheric temperature and humidity responses

436

437 Consistent with previous studies, the early ISM onset in the experiments is tightly linked
438 to the evolution of the tropospheric temperature over the Asian land mass, north of 20°N
439 (shading in Fig. 5; e.g. Li and Yanai 1996; and He et al. 2003; Xavier et al. 2007; Prodhomme et
440 al. 2015; Sabeerali and Ajayamohan 2017), and to the increase of the near surface humidity north
441 of the Equator (contours in Fig. 5; Goswami 2005; Ramesh Kumar et al. 2009; Goswami et al.
442 2010; Chakraborty and Agarwal 2017).

443 Substantial enhanced warming is observed at low-levels between 25°-35°N
444 corresponding to regions from northwest India extending toward Pakistan and part of
445 Afghanistan suggesting that the artificial decrease of the albedo in the two experiments helps in
446 enhancing the surface heat low over these regions during late spring and early summer. As a
447 further proof of this assertion, Figure 6 displays the spatial patterns of TS, 850-hPa wind and
448 MSLP anomalies during ISM in Desert_m20. As expected, the largest MSLP decreases occur

449 mainly over the hot subtropical desert, where TS increases mostly due to the albedo perturbation
450 introduced in Desert_m20 (Table 4). Interestingly, SST response over the IO in Desert_m20 is of
451 very modest amplitude despite of the fact that MSLP anomalies extend well over the ocean (Fig.
452 6). This indicates that TS anomaly over land is the major driver of the extended MSLP trough
453 and the associated low-level circulation, which develops during boreal summer in Desert_m20.
454 Similar features, but of reduced amplitude and spatial extension, are seen in Desert_Arab_m20
455 (not shown).

456 Tropospheric warming is also observed at nearly all pressure levels northward of 20°N in
457 the domain 55°-75°E in Desert_m20 during May-June, enhancing significantly the Meridional
458 Tropospheric Temperature Gradient (MTTG) in the Indian sector (Fig. 5a-f). On the other hand,
459 such changes in the upper atmosphere (above 500-hPa) are intriguingly much less pronounced in
460 Desert_Arab_m20 with a significant tropospheric warming mostly restricted to atmospheric
461 levels below 500-hPa (see Fig. 5g-l). These results from Desert_m20 suggest the key-role of the
462 hot subtropical deserts extending across Afro-Asian continents in the tropospheric temperature
463 and rainfall response over the Indian Sector. It also demonstrates that the regional differences in
464 the mid and upper troposphere are key factors controlling the differences in ISM responses
465 between Desert_m20 and Desert_Arab_m20 (Fig. 4b-c; see He et al. 2003; Dai et al. 2013).

466 In order to elucidate the origin of the tropospheric warming at higher levels during the
467 onset and withdrawal phases of ISM in the two perturbation experiments, Figure 7 shows the
468 anomalous 300-hPa wind and temperature responses in Desert_m20 and Desert_Arab_m20,
469 during May, June and September. Well defined anticyclonic circulation and warm temperature
470 anomalies are found at 300-hPa over the whole hot subtropical desert (the Arabia and Middle-
471 East deserts) during May-June in Desert_m20 (Desert_Arab_m20) illustrating the strong

472 baroclinicity of the atmospheric response (Figs. 7a-b and 7d-e; see Fig. 6 for the surface response
473 in Desert_m20) to the albedo perturbations. This suggests first a local contribution from surface
474 heating of the Sahara and Arabian deserts and the elevated regions of Iran, Afghanistan and
475 Pakistan.

476 Despite these similarities, the overall 300-hPa temperature response in Desert_m20 is
477 overwhelming compared to Desert_Arab_m20 during the ISM onset phase (see Figs. 7a-b and
478 7d-e). The most notable feature is the existence of a pronounced upper level warming (ranging
479 from 1 to 1.5°C) between 20° to 35°N and extending far away eastward in Desert_m20. This
480 suggests that the enhanced warming of the atmospheric column over the whole hot subtropical
481 desert in Desert_m20 (see Table 4) may also be advected eastward by the climatological upper-
482 level westerlies during April-June (figures not shown). These results are consistent with earlier
483 studies arguing that the heat advection associated with the Sub-Tropical westerly Jet stream
484 (STJ) and related planetary waves during April-June (Figure not shown) can influence the
485 temperature of the landmass lying north of ISM region (Xavier et al. 2007). Similar anomalous
486 features are observed during September in Desert_m20 (Fig. 7c). On the other hand, in
487 Desert_Arab_m20, the upper troposphere shows only subtle warming response during the ISM
488 onset and withdrawal phases (Figs. 7d,e and f).

489 The reasons for this upper-level differential warming between the two experiments are
490 further explored in Figure 8, which shows latitudinal vertical cross-sections of monthly
491 temperature (in shading) and vertical velocity (in contours) anomalies averaged over the North
492 African longitudes 20°W-30°E. In April-May, the lower troposphere gets heated from below
493 through the exchange of SH fluxes in Desert_m20 (Table 4) and this low-level warming expands
494 vertically in the middle and upper atmosphere (Fig. 8a-b). By June, upper-level (e.g. above 300-

495 hPa) warming starts intensifying around 20°-30°N, coinciding with the appearance of wide
496 spread upper-level subsidence northward of 25°N (Fig. 8c) which is manifested in adiabatic
497 descent causing warming in the upper atmospheric column. Note that warming may also be due
498 to the horizontal temperature advection as suggested by Cherchi et al. (2014) based on their
499 study over the same domain. Thereafter, this upper-level warming gets slightly weakened in
500 agreement with the disappearance of upper-level subsidence anomalies during July and August,
501 but both the anomalies (i.e. upper-level warming and subsidence) are again enhanced in
502 September. This remarkably concurs with the observed waning of enhanced ISM convection
503 during July and August, but its revival during September-October in Desert_m20 (Fig. 4b and
504 Fig. S2). In other words, this coupled pattern of upper-level warming and subsidence anomalies
505 at subtropical latitudes of North Africa is consistent with the monsoon-desert paradigm of
506 Rodwell and Hoskins (2001; see their Fig. 5) and is probably induced by the enhanced deep
507 convection over ISM region during both the first and last parts of boreal summer in Desert_m20
508 (see Fig. 4b and Fig S2). During July-August (Fig. 8d-e), one can see the emergence of large-
509 scale ascending motion and warming anomalies between 300 and 500-hPa from 15° to 30°N,
510 which coincide with the strengthening of the WAM in Desert_m20 (Fig. S2d-e).

511 In contrast, the tropospheric warming over Africa is quite feeble in Desert_Arab_m20
512 (right panels of Fig. 8), where the surface albedo is not perturbed over the Sahara region and
513 despite of the fact that ISM is also enhanced in this experiment. There appears to be only modest
514 descent and warming around 20°-35°N at upper-levels throughout June to September, which
515 probably coincides with the convection anomalies over ISM region (Rodwell and Hoskins 2001).
516 The modulation of the WAM and the associated warming in the middle atmosphere is also fully
517 missing here in contrast to Desert_m20. Finally, the differential response in the middle and upper

518 troposphere over Africa between the two experiments, further suggests the significant role of the
519 Sahara for sustaining the warming of the whole tropospheric column and the subsequent
520 eastward warm advection at upper-levels at the subtropical latitudes in Desert_m20 (Fig. 7).

521

522 4.3 ISM circulation response

523 We now examine the ISM atmospheric circulation response in Desert_m20, which shows
524 the larger ISM rainfall response.

525 In Desert_m20, the atmospheric wind anomalies observed in May (Fig. 9a) are
526 reminiscent of the mean-conditions observed during June in CTRL (not shown). One can see a
527 deep layer (between the surface and 600-hPa, Figs. 5b, 6a and 9a) of anomalous moist south-
528 westerlies blowing across the Arabian Sea extending up to 15°N, instead of a shallow layer
529 (between the surface and 800-hPa) of weak and drier low-level climatological westerlies
530 extending only 10°N (figure not shown). Consistently, Desert_m20 shows also the early
531 appearance of a pronounced Tropical Easterly Jet (TEJ, Wang 2006) in May, anomalously
532 extending northward between Equator and 20°N, along with the anomalous northward
533 displacement of STJ at upper-levels. The mean easterly shear is also enhanced north of the
534 equator in May, thus paving the way for an early ISM onset (Goswami 2005; Xiang and Wang
535 2013; Prodhomme et al. 2015; Parker et al. 2016). In other words, all these features confirm that
536 the stronger asymmetry of mean easterly shear and moisture distribution with respect to the
537 equator may be key-factors for explaining the early and abrupt monsoon transition in
538 Desert_m20 experiment (e.g. Webster and Yang 1992; Wang and Xie 1996; Xie and Wang 1996;
539 Jiang et al. 2004).

540 This also highlights a strong association between the seasonal march of the whole ISM
541 circulation and the heating over the hot subtropical desert since all the above described
542 atmospheric anomalies, especially the strengthened TEJ and the enhanced vertical easterly shear
543 observed between 10° - 25° N, are consistent with the thermal wind relationship and the warming
544 of the whole tropospheric column northward of 20° - 25° N observed from May to June (Fig. 5b-c;
545 e.g. Dai et al. 2013; Xavier et al. 2007; Ueda et al. 2006; Wu et al. 2012).

546 As a further confirmation of this tight thermal wind relationship for explaining the
547 anomalous ISM circulation in Desert_m20, the upper-level easterly and vertical shear anomalies
548 subsequently weaken during July-August (Fig. 9c-d) as the tropospheric warm anomalies
549 northward of 25° N in Desert_m20 also show similar signature (Figs. 5c-d). Such damping of the
550 atmospheric and rainfall anomalies during the peak phase of the monsoon in Desert_m20 may be
551 related to the dominance of the convective processes for regulating the evolution of the MTTG
552 and ISM circulation through latent heat release in the troposphere (Wang 2006; Xavier et al.
553 2007; Prodhomme et al. 2015) or to competitions between the WAM and ISM (see shading in
554 Fig. S2). Finally, the revival of the ISM circulation and rainfall during September (Figs. 4b and
555 9e) is again associated with an enhanced tropospheric warming in the northern parts of the
556 domain (Fig. 5f). This concomitant evolution of ISM rainfall and MTTG confirms again the
557 relevance of the MTTG to measure the ISM strength in our experiments.

558

559 **5. ISM response to increased surface albedo over the hot subtropical desert**

560 Here, we address the robustness and possible nonlinear nature of the relationship between
561 ISM and hot subtropical desert with the help of Desert_p20 and Desert_Arab_p20 experiments
562 (see Table 1 for more details). For conciseness, we will illustrate the responses mostly by using

563 Desert_p20, as Desert_Arab_p20 shows similar features, though of modest amplitude, as already
564 illustrated by Desert_Arab_m20 in Section 4.

565

566 5.1 Energy budget and ISM rainfall response

567 Despite of the fact that the amplitude of the artificial perturbations imposed in
568 Desert_p20 and Desert_m20 is exactly the same, the response in the simulated albedo and net
569 shortwave radiation budgets at the surface and TOA are stronger in Desert_p20 (Table 6) with a
570 decrease in solar shortwave radiation absorption at the ground (about -46 Wm^{-2}) and at TOA
571 (about -40 Wm^{-2}), as compared to those in Desert_m20 ($+31 \text{ Wm}^{-2}$ and $+25 \text{ Wm}^{-2}$, respectively at
572 surface and TOA, see Table 4). As less radiative energy is available at surface, both the surface
573 temperature and upward long-wave emission decrease substantially in Desert_p20, much more
574 than the corresponding increase for Desert_m20 (see Tables 4 and 6). However, this reduced
575 upward long-wave emission from the surface is fully compensated by a parallel decrease of the
576 downward long-wave emission since the net long-wave radiation budgets at surface are almost
577 similar in Desert_p20 and Desert_m20, with an excess of about 6 Wm^{-2} compared to CTRL. As a
578 consequence, the loss of shortwave radiation at the surface in Desert_p20 (e.g. -46 Wm^{-2}) is
579 almost exclusively compensated by a large decrease of the SH flux (e.g. 38 Wm^{-2}) since the LH
580 flux anomaly is very small (e.g. 2 Wm^{-2}). This highlights the key role of the SH flux anomalies
581 over the Desert for explaining the differences in the amplitude of ISM response between
582 Desert_p20 and Desert_m20. These asymmetrical results are similarly seen in the
583 Desert_Arab_p20 experiment as well (e.g. Tables 4 and 6).

584 This asymmetric nature of the atmospheric response is also observed over ISM region as
585 Table 7 shows decreased rainfall in the Desert_p20 (-3.2 mm day^{-1}) and Desert_Arab_p20 (-2.2

586 mm day⁻¹) experiments. Quantitatively, these rainfall responses in Desert_p20 and
587 Desert_Arab_p20 represent, respectively, a decrease of 53% and 36%, while the corresponding
588 increases are only of 34% and 16% in Desert_m20 and Desert_Arab_m20.

589 This decreased ISM rainfall in turn implies cloud-less conditions and less moisture,
590 which is further manifested by an important decrease in TOA albedo over the Indian domain (-
591 8%, see Table 7). Accordingly, there is stronger increase in net solar radiation both at TOA and
592 surface over the ISM region (e.g. 37 Wm⁻² at TOA and 45 Wm⁻² at surface in Desert_p20 for
593 example; see Table 7), with the net long-wave radiation partially offsetting it (e.g. -28 Wm⁻² at
594 TOA and -13 Wm⁻² at surface in Desert_p20; see Table 7), which is further consistent with a
595 strong surface warming (e.g. 2 and 1 K in Desert_p20 and Desert_Arab_p20, respectively; see
596 Table 7). Finally, as expected, the LH (SH) flux decreases (increases) over the ISM region
597 compared to CTRL since less moisture is available at the surface (see Table 7).

598 Figure 10 shows the seasonal evolution of rainfall response along 70°-90°E, from
599 Desert_p20 and Desert_Arab_p20 experiments. The response from Desert_p20 includes a slower
600 monsoon transition, a decrease in ISM rainfall during the whole boreal summer, which, in
601 conjunction with a late ISM onset and an early ISM withdrawal, results in a pronounced
602 shrinking of the length and intensity of the rainy season (Fig. 10a). Desert_Arab_p20 also shows
603 a similar response, but with a less obvious modulation in the length of the rainy season (Fig.
604 10b).

605 Based upon the responses in energy budget and rainfall (compare Figure 4 and Table 4
606 with Figure 10 and Table 6), one can infer that the positive perturbation experiments (e.g.
607 Desert_p20 and Desert_Arab_p20) consistently produce stronger responses than their negative
608 counterparts during the whole boreal summer season. This illustrates the key importance of the

609 fluctuation in the SH fluxes over the hot subtropical desert for the ISM rainfall and circulation as
610 shown next.

611

612 5.2 Tropospheric temperature and humidity responses

613 Figure 11 shows the seasonal evolution (from April to September) of the tropospheric
614 temperature and specific humidity responses from the Desert_p20 as latitude-pressure diagrams
615 averaged between 55°-75°E. Substantial tropospheric cooling is observed northward of 20°N,
616 which expands vertically in the middle and upper troposphere from April to June (Fig. 11a-c).
617 During July and August, this tropospheric cooling decreases slightly, especially in the middle
618 troposphere (e.g. between 600 and 400-hPa), but is again well pronounced in September (Fig.
619 11d-f). These tropospheric cooling pattern and evolution look symmetrically opposite to those
620 noticed in Desert_m20 (e.g. Fig. 5, see section 4), implying a reduced MTTG over the region,
621 especially at the start and end of ISM consistent with the ISM rainfall evolution in Desert_p20
622 (Fig. 10a). There is also a large reduction in meridional humidity gradient, which, in combination
623 with tropospheric cooling, may stabilize the atmosphere and hence explains the rainfall
624 suppression over the Arabian Sea and the Indian land mass during ISM.

625 We have also examined the seasonal evolution of temperature and vertical velocity
626 anomalies as latitudinal vertical cross-sections averaged over the longitudes 20°W-30°E in
627 Desert_p20 (figure not shown for brevity). In April-May, the lower troposphere (particularly
628 north of 20°N) gets asymmetrically cold from the surface (Table 6), which expands vertically
629 with time as in the Indian domain. Notably, the signature of cooling is prevalent throughout the
630 column, but with no amplified cooling at the upper-level, even though the vertical motion shows
631 anomalous tropospheric ascent north of 20°N. The implication is that the enhanced cooling,

632 which is prevalent throughout the troposphere, is of local origin due to the imposed artificial
633 surface albedo perturbation and the related enhanced SH flux anomaly over the deserts. This is
634 further illustrated by Fig. 12, which displays the MSLP, TS and 850-hPa anomalous patterns in
635 Desert_p20. A cold surface temperature anomaly coupled with an anomalous anticyclonic
636 response persists at the surface of the hot subtropical desert during the whole ISM. Interestingly,
637 as in Desert_m20 (Fig. 6), the SST response is only of small amplitude, even on the Arabian Sea.

638 Next we show the anomalous responses in 300-hPa wind and temperature for May, June
639 and September from Desert_p20 (Fig. 13) when the MTTG is largely reduced over the Indian
640 domain (e.g. Fig. 11b, c and f). Interestingly, at this level, one can find an anomalous cyclonic
641 circulation over the hot subtropical desert during May-June and September. This is in contrast to
642 the 300-hPa anomalous anticyclonic circulation over the same region in Desert_m20 (as noted in
643 Section 4), highlighting again the strong baroclinicity of the atmospheric response to the surface
644 albedo perturbations. Also, there is wide spread upper level cooling along the subtropical
645 latitudes of the hot subtropical desert region, which is quite analogous to the pronounced upper
646 level warming in Desert_m20. This may be again a signature of climatological upper-level
647 westerly wind acting on the temperature anomalies, following the interpretation of the eastward
648 temperature advection in Section 4.

649 On the other hand, the tropospheric cooling is much weaker over both Sahara and Indian
650 regions in Desert_Arab_p20 experiment (figures not shown), which suggests again the possible
651 role of the Sahara in explaining the differences between Desert_p20 and Desert_Arab_p20
652 (Table 6). The specific role of the Sahara in the experiments will be discussed in more details in
653 Section 6.

654

655 5.3 ISM circulation response

656 For more detailed understanding on the underlying processes, we next examine the ISM
657 circulation in Desert_p20 (Fig. 14). In Desert_p20, one can see dry low-level (between the
658 surface and 800-hPa) easterly anomalies over the Arabian Sea in May (Figs. 11b, 12a and 14a),
659 in contrast to deep anomalous moist westerly anomalies in Desert_m20 (Figs. 5b, 6a and 9a).
660 These surface easterly anomalies in Desert_p20 are part of the anomalous surface anticyclone,
661 which forms with the anomalous cyclonic circulation at higher levels (from 600 to 300-hPa), a
662 robust baroclinic structure to the west of ISM region during the whole boreal summer (Figs. 12
663 and 13). This is further supported by tropospheric descent anomalies north of 20°N over the
664 Indian landmass, which expand considerably to the south (up to 10-15°N) and vertically from
665 May to June (not shown), thus impeding the ISM onset progression and hence leading to a
666 delayed onset and weaker ISM in Desert_p20. Accordingly, the appearance of TEJ and
667 northward migration of STJ at upper levels are conspicuously absent in Desert_p20 during May
668 and June (Fig. 14a-b). These aforementioned features further confirm that the stronger MTTG
669 along with the equatorial asymmetry of mean easterly shear and moisture are vital in determining
670 the seasonal transition and maintenance of the ISM system.

671

672 5.4 Origin of the asymmetry in ISM rainfall responses between Desert_p20 and Desert_m20

673 A relevant question here is on the origin of the enhanced large-scale rainfall suppression
674 and dryness over the Indian land mass longitudes in Desert_p20 (Fig. 10) compared to the ISM
675 rainfall increase in Desert_m20 (Fig. 4).

676 Here we carried out additional diagnostics using a moisture budget analysis to discern the
677 underlying moistening and drying process and to unravel the origin of non-linear anomalous

678 responses between the two simulations (Desert_m20 and Desert_p20). The time mean budget
679 moisture budget equation (e.g. Dixit et al. 2018) is as follow:

$$680 \quad P = E - u \frac{\partial q}{\partial x} - v \frac{\partial q}{\partial y} - w \frac{\partial q}{\partial p} \quad (1)$$

681 Here P is rainfall and E is evaporative latent heat flux over the ISM domain (see Tables 5
682 and 7). Also, u, v and w are zonal, meridional and vertical pressure velocity, with q being
683 specific humidity. The equation can be applied to the control (CTRL) and perturbed (e.g.
684 Desert_m20 and Desert_p20) simulations to get the anomalous responses (against CTRL).

685 Figure 15 shows the vertical profile of anomalous moisture advection tendencies for the
686 two experiments (Desert_m20, left panels and Desert_p20, right panels). In Desert_m20 (Fig
687 15a-e, see the shading), the anomalous moistening due to vertical advection tendency prevails
688 throughout the middle troposphere during the early and late parts of ISM and is in phase with the
689 seasonal rainfall evolution depicted in Figure 4b, and hence dominantly contributes to the
690 anomalous rainfall evolution. Horizontal anomalous advection in Desert_m20 (contours in
691 Figure 15a-e) moistens the lower boundary layer due to the early penetration of monsoon flow to
692 Indian latitudes in the Arabian Sea (see Fig 6a) and thus also contributes to the rainfall
693 anomalies, but only during the first phase of ISM in Desert_m20. It seems that the horizontal
694 advection also moistens the lower-troposphere to the northward edge of the rainfall peak, thus
695 paving the way for the seasonal migration of rainfall and hence the abrupt monsoon transition in
696 Desert_m20.

697 In Desert_p20 (Fig 15f-j), the budget responses are of amplified magnitude throughout
698 the season relative to Desert_m20 and for both the vertical and horizontal moist advection terms,
699 thus highlighting the non-linear nature of the anomalous dynamical response. This results in a

700 pronounced and non-linear tropospheric drying effect and the consequent large-scale rainfall
701 suppression over the Indian continent as described above (e.g. Section 5.2 and Fig. 10).

702 In other words, the asymmetric responses in ISM circulation and rainfall between
703 Desert_p20 and Desert_m20, as discussed in Section 5.2, are tightly associated with the non-
704 linearity in advection moisture tendencies. This further points towards some important positive
705 feedbacks between the moisture convergence and the ascending motion (figure not shown) in the
706 atmospheric column over land, which are much more active in Desert_p20, thus contributing to
707 the large-scale suppression of rainfall over Indian continent in this simulation (as depicted in Fig.
708 10).

709 For explaining these differences one can further envisage the enhanced responses in TS
710 and MSLP in Desert_p20 (see Table 6 and Fig. 12b-e) as compared to Desert_m20 (Table 4 and
711 Fig. 6b-e), which thus contribute to the differences between Desert_p20 and Desert_m20 through
712 geostrophic low-level wind adjustments and related anomalous moisture convergence at the low-
713 levels during ISM (Figs. 6 and 12; see also Samson et al. 2016; Terray et al. 2017).

714 In a similar manner, the decrease of the LH flux over the Indian landmass in Desert_p20
715 (e.g. 24 Wm^{-2} ; see Table 7) is stronger in amplitude than its increase in Desert_m20 (e.g. -10
716 Wm^{-2} ; see Table 5). This reduced evaporation over land can again reduce the local moisture
717 availability and, thus, contributes to further reduce the local diabatic heating of the atmospheric
718 column and feedbacks on itself, leading to a stronger ISM response in Desert_p20. This stronger
719 local effect in Desert_p20 may also reflect the tendency of the ground over India to become more
720 easily saturated in Desert_m20, thus limiting the local LH flux and diabatic heating anomalies in
721 this simulation. Recently, Halder et al. (2016) and Paul et al. (2016), using sensitivity
722 experiments with regional climate models, similarly showed that the decrease in local moisture

723 can efficiently reduce the low-level moist static energy, thereby increasing the atmospheric
724 stability and suppressing the large-scale convective instability as well as convective activity.

725 It should be reminded, however, that in our simulations the activation of these positive
726 feedbacks, involving local evaporation, moisture convergence and diabatic heating over India, is
727 ultimately linked to the remote radiative forcing induced by the surface positive albedo
728 perturbation over the hot subtropical desert to the west of the Indian domain in Desert_p20.

729

730 **6. Discussion**

731

732 6.a Specific role of the Sahara in the ISM circulation and rainfall responses

733 In earlier sections, we infer that the inclusion of the Sahara made a significant difference
734 in radiative forcing, but also in the ISM rainfall and circulation responses based on the
735 comparison of our Desert and Desert_Arab simulations. While these results suggest the potential
736 significance of the Sahara, these simulations are not sufficient to delineate properly the specific
737 role of the Sahara in the ISM response because the magnitude of radiative forcing in both
738 simulations is hugely different and the response is probably nonlinear as highlighted in the
739 previous sections.

740 In order to bring out the relative importance of Sahara and the Arabian-Pakistan-Thar
741 deserts for ISM, we carried out additional perturbation runs only over the Sahara region (e.g.
742 with artificial changes of -0.2 and 0.2 in the background land albedo over the Sahara region; see
743 Table 1 and Fig. 1b for details).

744 The ISM rainfall response in these experiments, which we refer as Desert_Sahara_m20
745 and Desert_Sahara_p20, is displayed in Fig. 16. The results show a significant response with

746 stronger (weaker) amplitude for a negative (positive) albedo perturbation as compared to those
747 from Desert_Arab_m20 (Desert_Arab_p20) experiments (Figs. 4c, 10b). This demonstrates first
748 the specific role of the Sahara warming/cooling on the ISM. The response is concentrated during
749 the onset and first phases of the ISM, thus indicating a short-lived influence with no significant
750 modulation in the length of the ISM rainy season as in the Desert_Arab experiments.

751 Again, the amplitude of the ISM response is stronger with a positive albedo perturbation
752 (Fig. 16b), further highlighting the nonlinearity of the governing processes. Furthermore, the
753 response to the forcing of the whole NH subtropical desert (i.e. ISM response from Desert_m20
754 and Desert_p20 experiments; see Figs. 4b, 10a) is significantly stronger than the addition of
755 those responses due to the Sahara (Figs. 16) and Arabia-Middle-East (Figs. 4c and 10b) deserts
756 separately, thus further attesting to the non-linearity of the relationship between the NH
757 subtropical deserts and ISM.

758

759 6.b Robustness of the results

760 Our aforementioned model results however need obviously some cautionary remarks,
761 because they can be model dependent. Furthermore, the specific coupled model SST biases may
762 also affect the realism of the findings. Thus, the robustness of these results/conclusions needs to
763 be verified with other coupled/forced models, especially the asymmetric ISM response. Hence in
764 this direction, we have performed the same experiments with another coupled model, the CFSv2
765 model, currently in use at NCEP and IITM for seasonal forecasting (see Saha et al. 2014 and
766 Terray et al. 2017 for details). Importantly, CFSv2 includes completely different atmospheric,
767 oceanic and land dynamic components than SINTEX. Another interesting point is that both
768 CFSv2 and SINTEX share exactly similar rainfall biases, however their SST biases are nearly

769 opposite in most of the tropical regions (Terray et al. 2017). More specifically, CFS has a cold
770 SST bias over the Indian Ocean (Swapna et al. 2015), while SINTEX has a warm bias over the
771 Indian Ocean (Joseph et al 2012).

772 Figure S3 shows the ISM rainfall seasonal cycle in the control simulation and the
773 responses in various experiments performed with CFSv2. A comparison of Figs. 4, 10 and S3
774 shows that the ISM rainfall responses are qualitatively and quantitatively very similar in the two
775 models despite their important differences in parameterizations or SST bias. More specifically,
776 there is an abrupt and early ISM onset transition in the CFSv2 Desert_m20 experiment, with a
777 significant increase of ISM length (Fig. S3b). Figure S4 then highlights the strong association
778 between the seasonal march of the ISM circulation and heating over the hot subtropical desert
779 with associated increase of the inter-hemispheric MSLP gradient over western part of Indian
780 domain, which is further instrumental in enhanced cross-equatorial flow and moisture
781 convergence to the Indian landmass. The circulation and heating patterns in CFSv2 are very
782 similar to those obtained with SINTEX (see Fig. 6).

783 Again consistent with SINTEX, qualitatively opposite rainfall response and associated
784 heating patterns, but of asymmetric nature (e.g. stronger), are produced in Desert_p20 (Figs. S3c
785 and S5), with an anomalous coupled pattern of cold surface temperature and anticyclonic
786 circulation lying over the hot subtropical desert. Furthermore, the Desert_Arab_m20 and
787 Desert_Arab_p20 experiments confirm the short-lived ISM response (compare Figs. 4c, 10b and
788 S3d-e), indicating again the prominent role of the hot subtropical desert extending across the
789 Afro-Asian continent on the ISM response. In other words, the nonlinearity of the ISM response
790 is also found in the CFSv2 experiments.

791

792 **7. Conclusion and Perspectives**

793 During boreal summer, the South Asian monsoon and subtropical deserts coexist at the
794 same latitudes over the African-Asian continent. Previous studies have tried to explain this
795 coexistence, as well as the mutual relationships between these two very different climates (Yang
796 et al. 1992; Webster 1994; Rodwell and Hoskins 1996, 2001; Wu et al. 2009; Bollasina and
797 Nigam 2011a; Tyrlis et al. 2013; Cherchi et al. 2014; Krishnamurti et al. 2010).

798 However, the significant relationship between monsoon rainfall biases and surface
799 temperature, albedo and humidity biases over the adjacent subtropical deserts in current state-of-
800 the-art climate models (Samson et al. 2016; Agrawal and Chakraborty 2016; Terray et al. 2017)
801 warrants further investigations on the possible role of the NH subtropical deserts on ISM
802 intensity and evolution. Moreover, desert amplification over the NH subtropical deserts during
803 boreal summer is one of the main modes of surface temperature warming associated with
804 anthropogenic climate change (Zhou 2016; Wei et al. 2017).

805 These considerations motivate our re-examination of the ISM-desert paradigm. In order
806 to unravel new facets on the role of the hot subtropical desert on ISM, the present study uses a
807 global ocean-atmosphere coupled model and a set of carefully designed experiments by
808 perturbing surface albedo in different ways over the hot subtropical desert (or parts of it) to the
809 west of the South Asian domain.

810 The analysis of up-to-date satellite datasets, atmospheric re-analyses and our control
811 coupled simulation suggests that the model broadly reproduces the radiation budgets as seen in
812 re-analyses and observed datasets. However there are large uncertainties in the surface and TOA
813 radiation budgets over the NH subtropical desert region (Fig S1), with the budget estimates
814 showing inconsistency regionally. While the budget estimates from reanalysis (ERAi) show a

815 radiative excess (despite of large upward long wave emission) throughout the NH desert region
816 during boreal summer, in contrast, our model estimates have rather neutral radiation budget over
817 the Sahara Desert. A few early studies (Smith 1986a,b; Vinoj et al. 2014) already suggested that
818 the hot subtropical desert (e.g. over Arabian Peninsula) do not show the radiative sink property
819 during boreal summer which is again in contrast with the observed estimates from CERES-
820 EBAF (Fig. S1a).

821 Our desert albedo perturbation experiments further confirm that variations of the surface
822 radiation budget over the Desert or parts of it, may impact the boreal summer monsoon far away
823 since ISM evolution and intensity are affected with opposite polarity to prescribed negative and
824 positive albedo perturbations over the whole hot subtropical desert or part of it (e.g. the Arabia
825 and Middle-East or Sahara deserts). Interestingly, the Sahara is also playing a significant role in
826 the strength of this ISM response and not only the Arabian-Iran-Thar desert adjacent to the ISM
827 domain, as demonstrated by the comparison of our Desert, Desert_Arab and Desert_Sahara
828 simulations (Figs 4, 10 and 16).

829 A second (expected) finding is that by darkening the surface of the deserts and arid
830 regions to the west of the South Asian domain, radiative heating of the surface is increased.
831 Interestingly, this additional heating is not solely due to the surface's albedo direct effect on solar
832 absorption, but also to an increase of the net long-wave radiation budget. This suggests the
833 existence of a positive feedback amplifying the original surface warming due solely from the
834 surface albedo perturbation. The modest absolute changes of latent heating and rainfall over the
835 deserts in the reduced albedo experiments seem to be sufficient to amplify this original surface
836 warming by an even larger increase of the downward long-wave emission received at the

837 surface. This leads ultimately to more excess in the surface and TOA net radiation budgets over
838 the deserts (e.g. Table 4).

839 This heating over the desert is communicated to the overlying atmosphere mainly by
840 additional SH flux and dry convection due to the very low ambient soil moisture. It generates a
841 well-defined baroclinic atmospheric response with cyclonic wind and negative MSLP anomalies
842 near the surface (e.g. Fig. 6), and anticyclonic wind anomalies aloft in the middle and upper
843 troposphere (Fig. 7). Such tropospheric heating and atmospheric responses affect the adjacent
844 monsoon systems, including ISM, in several co-operating ways.

845 First, it is found that the tropospheric warming over the hot subtropical desert results also
846 in warmer upper and middle troposphere over the northern latitudes of the Indian subcontinent
847 due to warm advection by the STJ in the upper troposphere during the onset and withdrawal
848 phases of ISM (Figs. 5-8). This leads to an increased MTTG over the South Asian domain,
849 which favors a strengthening of the easterly vertical wind shear (e.g. in agreement with thermal
850 wind balance; Fig. 9) during the first and last phases of ISM, in addition to the baroclinic
851 atmospheric structure directly linked to the albedo perturbation described above.

852 Second, the enhanced surface heat low over the Arabian Peninsula and the Iran-Pakistan-
853 Thar desert (due to the local surface warming) increases significantly the inter-hemispheric
854 MSLP gradient over the western part of the Indian domain because this MSLP gradient is mainly
855 controlled by the seasonal evolution of the MSLP fields over the land (Fig. 6; Li and Yanai,
856 1996). This generates an enhanced cross-equatorial flow along the African coast, which
857 gradually becomes westerly over the Arabian Sea due to the Coriolis effect, enhancing the
858 moisture convergence over the Indian landmass during boreal summer (Samson et al. 2016;
859 Terray et al. 2017).

860 All these factors (e.g. the easterly vertical shear of the zonal wind and the additional
861 moisture north of the equator) contribute to explain why the length of the ISM is substantially
862 increased in our Desert_m20 simulation (Fig. 4b), by both an early onset and a late withdrawal of
863 the monsoon. Thus, our reduced albedo experiments further validate the well-known results that
864 a stronger MTTG along with stronger equatorial asymmetry of mean easterly vertical shear and
865 moisture distribution, are key-factors for explaining the early and abrupt ISM onset (e.g.
866 Goswami et al. 2010; Parker et al. 2016).

867 It must be noted, however, that this increase of the length of the rainy season is much
868 stronger in Desert_m20 than in Desert_Arab_m20 (Fig. 4b-c), highlighting the key-role of the
869 whole hot subtropical Afro-Asian desert in the response. The short-lived ISM response in
870 Desert_Arab_m20 corroborates previous studies, which show a significant association between
871 the modulations of the heat low spread across the northwest India, Pakistan, Afghanistan and
872 Iran regions during the pre-monsoon period with the ISM intensity in its initial period, but with
873 the land-atmosphere coupling losing its signature once the monsoon matures (Ramage 1966;
874 Mooley and Paolino 1988; Rai et al. 2015; Agrawal and Chakraborty 2016).

875 Qualitatively opposite results are obtained with positive albedo perturbations (e.g. the
876 Desert_p20 and Desert_Arab_p20 simulations). But, both the amplitude of the SH flux
877 anomalies over the deserts (Tables 4 and 6) and of the ISM response (Figs. 4 and 10) are
878 significantly stronger, highlighting non-linear characteristics of this response. In the positive
879 albedo perturbations runs, there is a larger reduction of the MTTG and the meridional humidity
880 gradient during the whole ISM season. Furthermore, the asymmetric responses in ISM
881 circulation and rainfall between Desert_p20 and Desert_m20 are tightly associated with the non-
882 linearity in advection moisture tendencies, as revealed through the application of moisture

883 budget analysis (see Section 5.2; Fig. 15). This points toward some important positive feedbacks
884 between the moisture convergence and the ascending motion in the atmospheric column over
885 land (Halder et al. 2016; Paul et al. 2016), which are much more active in this simulation than in
886 Desert_m20 despite of the fact that the amplitude of the imposed albedo perturbation is the same
887 in these simulations.

888 We also demonstrate the robustness of the land-surface heating over subtropical deserts
889 to the seasonal march of ISM by performing similar experiments with another independent
890 coupled model (see Section 6b). The results validate all our conclusions about the significant and
891 differential roles of the NH subtropical deserts, including the Sahara, on the ISM system since
892 they are independent of parameterization schemes used in the coupled models or SST biases
893 affecting the coupled models.

894 Finally, the robustness of the relationship between the ISM and the neighboring deserts to
895 its west found here suggests that these driest regions may play an increasingly significant role in
896 the ISM evolution in the future climate through the modulation of their intense surface warming.
897

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1132

1133 **Figure captions:**

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1135 **Figure 1:** Spatio-temporal evolution of rainfall climatology (mm day^{-1}) for May to September.

1136 The overlying contours represent rainfall estimates from GPCP, while shading is the same from

1137 CTRL (control simulation). The contours are drawn following the shading convention and the 5

1138 mm day^{-1} contour is highlighted in thick black colour. The pattern correlation between CTRL

1139 and GPCP for the tropical domain (i.e. within 20°S to 35°N) is indicated in each panel. The

1140 boxed area in (a), (b) and (d) represent the hot subtropical deserts regions where artificial albedo

1141 perturbation experiments are carried out (see Table 1 for more details).

1142 **Figure 2:** Seasonal evolution (May to September) of vertical component of velocity (shading,

1143 units in $10^{-2} \text{ Pa s}^{-1}$) and horizontal divergence (contours, units in 10^{-6} s^{-1}), along a pressure-

1144 longitude plane averaged over 15° - 30°N for ERAi (left panel, a-e) and CTRL (right panel, f-j).

1145 The negative (dashed) and positive (continuous) contours correspond, respectively, to absolute

1146 magnitudes of **1, 2, 3 and 4** units. The zero contours are highlighted in thick black colour.

1147 **Figure 3 (a-j):** Same Figure 2, but for the vertical component of velocity (shading, units in 10^{-2}

1148 Pa s^{-1}) and horizontal divergence (contours, units in 10^{-6} s^{-1}), averaged over 30° - 40°N . The

1149 negative (dashed) and positive (continuous) contours correspond, respectively, to absolute

1150 magnitudes of 1, 2, 3 and 4 units. The zero contours are highlighted in thick black colour.

1151 **Figure 4:** (a) Time-latitude evolution of rainfall climatology (mm day^{-1} , containing both land

1152 and oceanic points) averaged along 70° - 90°E , from CTRL. (b) and (c) time-latitude evolution of

1153 anomalous rainfall response (computed against the CTRL rainfall climatology, see Section 2)

1154 averaged along 70° - 90°E from Desert_m20 and from Desert_Arab_m20, respectively. In (b) and

1155 (c), the responses that are above the 95% confidence level according to a permutation procedure

1156 with 9999 shuffles are encircled.

1157 **Figure 5:** (Left Panel, a-f) Seasonal evolution (April to September) of response in temperature
1158 (shading, °C) and specific humidity (contours, $\times 10^{-3}$ Kg Kg⁻¹) along a pressure-latitude plane
1159 over the longitudes 55°-75°E, for Desert_m20 experiment. Right panel (g-l), same as (a-f), but
1160 for Desert_Arab_m20 experiment. The anomalous responses are computed against the CTRL
1161 climatology (see Section 2). The negative (dashed) and positive (continuous) contours
1162 correspond, respectively, to absolute magnitudes of 0.25, 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero
1163 contours are highlighted in thick black colour.

1164 **Figure 6:** Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS
1165 (shading, °C) and 850-hPa wind (ms⁻¹) from Desert_m20 experiment. Negative (dashed) and
1166 positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6
1167 units. The zero contours are not drawn. The anomalous response is computed against the CTRL
1168 climatology (see Section 2). Wind vectors of magnitudes exceeding 1 ms⁻¹ are only drawn.

1169 **Figure 7:** Map of the anomalous responses in wind and temperature at 300-hPa for (a) May, (b)
1170 June and (c) September, from Desert_m20 experiments. (d)-(f), same as (a)-(c), but from
1171 Desert_Arab_m20. The anomalous response is against the CTRL climatology (see Section 2).

1172 **Figure 8:** (Left Panel, a-f) Seasonal evolution (April to September) of response in temperature
1173 (shading, °C) and vertical component of velocity (contours, units is 10^{-2} Pa s⁻¹) along a pressure-
1174 latitude plane averaged over the longitudes 20°W-30°E, for Desert_m20 experiment. Right panel
1175 (g-l), same as (a-f), but for Desert_Arab_m20 experiment. The anomalous response is computed
1176 against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous)
1177 contours correspond respectively to magnitudes of 0.25, 0.5, 1, 2, 3, 4 and 5 units. The zero
1178 contours are highlighted in thick black colour.

1179 **Figure 9:** Seasonal evolution (May to September) of response in horizontal component of wind
1180 (zonal wind in shading and meridional wind in contours, ms^{-1}) along a pressure-latitude plane
1181 averaged over the longitudes 55° - 75° E, for Desert_m20 experiment. The anomalous response is
1182 computed against the CTRL climatology (see Section 2). The negative (dashed) and positive
1183 (continuous) contours correspond respectively to magnitudes of 0.5, 1, 1.5, 2, 2.5 and 3 units.
1184 The zero contours are highlighted in thick black colour.

1185 **Figure 10:** Time-latitude evolution of rainfall response (mm day^{-1} , averaged along 70° - 90° E
1186 from the sensitivity experiments. In (a) from Desert_p20 and in (b) from Desert_Arab_p20. The
1187 anomalous responses are computed against the CTRL climatology (see Section 2). In all panels,
1188 the responses that are above the 95% confidence level according to a permutation procedure with
1189 9999 shuffles are encircled.

1190 **Figure 11:** Seasonal evolution (April to September) of response in temperature (shading, $^{\circ}\text{C}$)
1191 and specific humidity (contours, $\times 10^{-3} \text{ Kg Kg}^{-1}$) along a pressure-latitude plane over the
1192 longitudes 55° - 75° E, for Desert_p20 experiment. The anomalous response is computed against
1193 the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours
1194 correspond, respectively, to magnitudes of 0.25, 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero contours
1195 are highlighted in thick black colour.

1196 **Figure 12:** Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS
1197 (shading, $^{\circ}\text{C}$) and 850-hPa wind (ms^{-1}) from Desert_p20 experiment. Negative (dashed) and
1198 positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6
1199 units. The zero contours are not drawn. The anomalous response is computed against the CTRL
1200 climatology (see Section 2). Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn.

1201 **Figure 13:** (a) Anomalous response in wind (ms^{-1}) and temperature ($^{\circ}\text{C}$, shading) at 300-hPa,
1202 during May from Desert_p20 experiment. (b) and (c), same as (a), but for June and September,
1203 respectively. The anomalous response is computed against the CTRL climatology (see Section
1204 2).

1205 **Figure 14:** Seasonal evolution (May to September) of response in horizontal component of wind
1206 (zonal wind in shading and meridional wind in contours, ms^{-1}) along a pressure-latitude plane
1207 averaged over the longitudes 55° - 75°E , for Desert_p20 experiment. The anomalous response is
1208 computed against the CTRL climatology (see Section 2). The negative (dashed) and positive
1209 (continuous) contours correspond, respectively, to magnitudes of 0.5, 1, 1.5, 2, 2.5 and 3 units.
1210 The zero contours are highlighted in thick black colour.

1211 **Figure 15:** (Left Panel, a-e) Seasonal evolution (May to September) of response in vertical
1212 moisture advection (shading, units is $10^{-3} \text{ Kg Kg}^{-1} \text{ day}^{-1}$) and horizontal moisture advection
1213 (contours, $\times 10^{-3} \text{ Kg Kg}^{-1} \text{ day}^{-1}$) along a pressure-latitude plane over the longitudes 75° - 90°E , for
1214 Desert_m20 experiment. Right panel (f-j), same as (a-e), but for Desert_p20 experiment. The
1215 anomalous response is computed against the CTRL climatology (see Section 2). The negative
1216 (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.1, 0.2,
1217 0.3, 0.4, 0.5 and 0.6 units. The zero contours are highlighted in thick black colour.

1218 **Figure 16:** (a) Time-latitude evolution of rainfall response (mm day^{-1} , averaged along 70° - 90°E
1219 from the Desert_Sahara_m20 experiment. In (b), same as (a) but for Desert_Sahara_p20
1220 experiment. The anomalous responses are computed against the CTRL climatology (see Section
1221 2). In (a) and (b), the responses that are above the 95% confidence level according to a
1222 permutation procedure with 9999 shuffles are encircled.

1223

1224 **Figure captions (Supplementary Figures):**

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1226 **Figure S1:** Climatological map of net radiation budget at TOA (Wm^{-2} , see Table 2 and Section 2
1227 for more details on radiation budget) for June to September period. In (a). CERES-EBAF, (b)
1228 ERAi and (c) CTRL

1229 **Figure S2:** Spatio-temporal evolution of rainfall (May to September, mm day^{-1}) response from
1230 the Desert_m20 experiment (in shading). The overlying contours are the rainfall response from
1231 Desert_Arab_m20. Negative (dashed) and positive (continuous) contours are drawn following
1232 the shading convention. The anomalous responses are computed against the CTRL climatology
1233 (see Section 2).

1234 **Figure S3:** In (a) to (c), same as that of Figure 4, but for the various experiments (CTRL,
1235 Desert_m20 and Desert_p20), conducted using CFSv2 coupled model (see Terray et al. 2017).
1236 Similarly, in (d) and (e), same as that of Figure 11, but for Desert_Arab_m20 and
1237 Desert_Arab_p20 simulations, using CFSv2. The length of integration (i.e. using CFSv2) in
1238 CTRL simulation is for 40 years, while it is 20 years for the Desert_m20 and Desert_p20. Note
1239 that the configuration used for the experiments also takes advantage of up-to-date satellite
1240 MODIS data for estimating the background snow-free albedo as described in Terray et al.
1241 (2017). The first 10 years of all the simulations are excluded for all the analyses presented in the
1242 text. In (b) and (c), the responses that are above the 95% confidence level according to a
1243 permutation procedure with 9999 shuffles are encircled.

1244 **Figure S4:** Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS
1245 (shading, $^{\circ}\text{C}$) and 850-hPa wind (ms^{-1}) from Desert_m20 experiment done using CFSv2 model.
1246 Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of
1247 0.5, 1.5, 3, 4 and 6 units. The anomalous response is computed against the CTRL climatology.

1248 Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn. For more details of the CFS runs,
1249 please refer to the legend of Fig. S3.

1250 **Figure S5:** Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS
1251 (shading, $^{\circ}\text{C}$) and 850-hPa wind (ms^{-1}) from Desert_p20 experiment done using CFSv2 model.
1252 Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of
1253 0.5, 1.5, 3, 4 and 6 units. The anomalous response is computed against the CTRL climatology.

1254 Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn. For more details of the CFS runs,
1255 please refer to the legend of Fig. S3.

1256

1257 **Table captions:**

1258 **Table 1:** Details of coupled model experiments as performed using SINTEX-F2 coupled model.

1259 The first 10 years of all the simulations are excluded for all the analyses presented in the text.

1260 See Section 2 for more details of the experiments.

1261 **Table 2:** The radiative and turbulent fluxes, at surface (SURF) and top of the atmosphere (TOA),

1262 averaged over the land points in the domain 15°-40°N and 20°W-75°E (e.g. hot subtropical

1263 desert region) for June to September (JJAS) season. Here, SW and LW are shortwave and long-

1264 wave radiations, respectively. Net Rad means net radiation following the formulation in Su and

1265 Neelin (2002). SH and LH are sensible and latent heat fluxes. TS and PR stands for surface skin

1266 temperature and rainfall, respectively. Fluxes are signed in the direction of the fluxes (i.e.

1267 negative sign indicates upward fluxes and vice-versa). Asterisk (*) in the Table represents that

1268 the rainfall (PR) is averaged using the GPCP data and not using ERAi estimates. All fluxes have

1269 energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and mm day^{-1} ,

1270 respectively.

1271 **Table 3:** The radiative and turbulent fluxes averaged over Indian landmass (land points only:

1272 75°-90°E, 5°-25°N) for JJAS season. Fluxes are signed in the direction of the fluxes (i.e.

1273 negative sign indicates upward fluxes and vice-versa). Asterisk (*) in Table represents that the

1274 rainfall (PR) is averaged using the GPCP data and not using ERAi estimates. All fluxes have

1275 energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and mm day^{-1} ,

1276 respectively.

1277 **Table 4:** The radiative and turbulent fluxes averaged over the land points in the hot subtropical

1278 desert domain (15°-40°N and 20°W-75°E) during JJAS season for the negative albedo

1279 perturbation experiments. Asterisk (*) in Table indicates that the flux values are averaged only

1280 over 15°-40°N and 35°-75°E (e.g. in the Desert_Arab_m20 experiment). Fluxes are signed in the
1281 direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). The upper part
1282 of the Table provides the climatological estimates from the sensitivity experiments while the
1283 lower part shows the anomalous responses (deviations from the CTRL climatology, see Section
1284 2). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K
1285 and mm day^{-1} , respectively.

1286 **Table 5:** The radiative and turbulent fluxes averaged over Indian landmass (land points only:
1287 75°-90°E, 5°-25°N) during JJAS season for the negative albedo perturbation experiments. Fluxes
1288 are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-
1289 versa). The upper part of the Table provides the climatological estimates from the sensitivity
1290 experiments while the lower part shows the anomalous responses (deviations from the CTRL
1291 climatology). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the
1292 units of K and mm day^{-1} , respectively.

1293 **Table 6:** The radiative and turbulent fluxes averaged over the hot subtropical desert domain
1294 during JJAS season for the positive albedo perturbation experiments. Asterisk (*) in Table
1295 indicates that the flux values are averaged only over 15°-40°N and 35°-75°E (e.g. in the
1296 Desert_Arab_p20 experiment). Fluxes are signed in the direction of the fluxes (i.e. negative sign
1297 indicates upward fluxes and vice-versa). The upper part of the Table provides the climatological
1298 estimates from the sensitivity experiments while the lower part shows the anomalous responses
1299 (deviations from the CTRL climatology). All fluxes have energy units (Wm^{-2}) and albedo is in
1300 %, while TS and PR have the units of K and mm day^{-1} , respectively.

1301 **Table 7:** The radiative and turbulent fluxes averaged over Indian landmass (land points only:
1302 75°-90°E, 5°-25°N) during JJAS season for the positive albedo perturbation experiments. Fluxes

1303 are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-
1304 versa). The upper part of the Table provides the climatological estimates from the sensitivity
1305 experiments while the lower part shows the anomalous responses (deviations from the CTRL
1306 climatology, see Section 2). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS
1307 and PR have the units of K and mm day^{-1} , respectively.

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Figures:

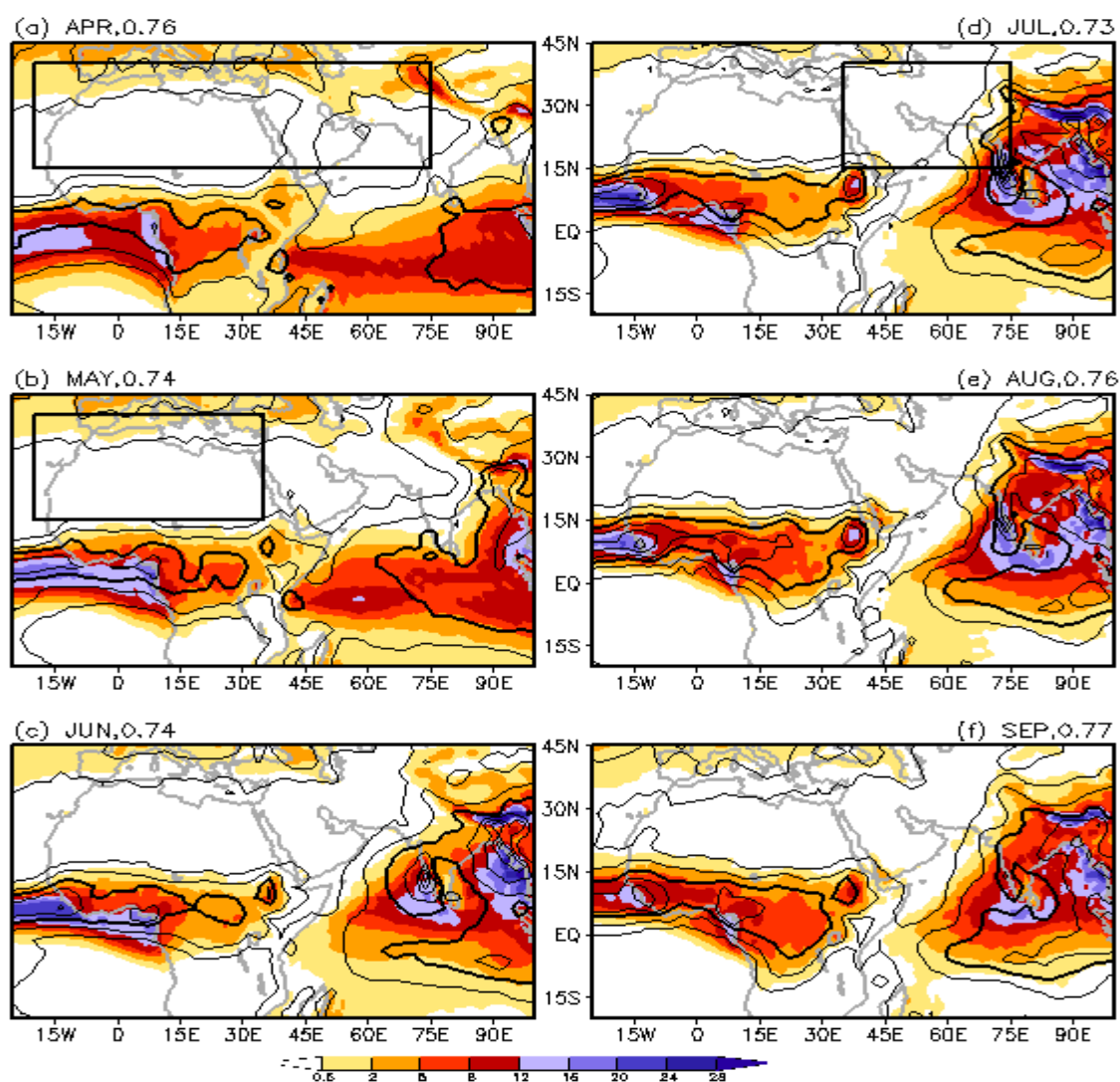


Figure 1: Spatio-temporal evolution of rainfall climatology (mm day^{-1}) for May to September. The overlying contours represent rainfall estimates from GPCP, while shading is the same from CTRL (control simulation). The contours are drawn following the shading convention and the 5 mm day^{-1} contour is highlighted in thick black colour. The pattern correlation between CTRL and GPCP for the tropical domain (i.e. within 20°S to 35°N) is indicated in each panel. The boxed area in (a), (b) and (d) represent the hot subtropical deserts regions where artificial albedo perturbation experiments are carried out (see Table 1 for more details).

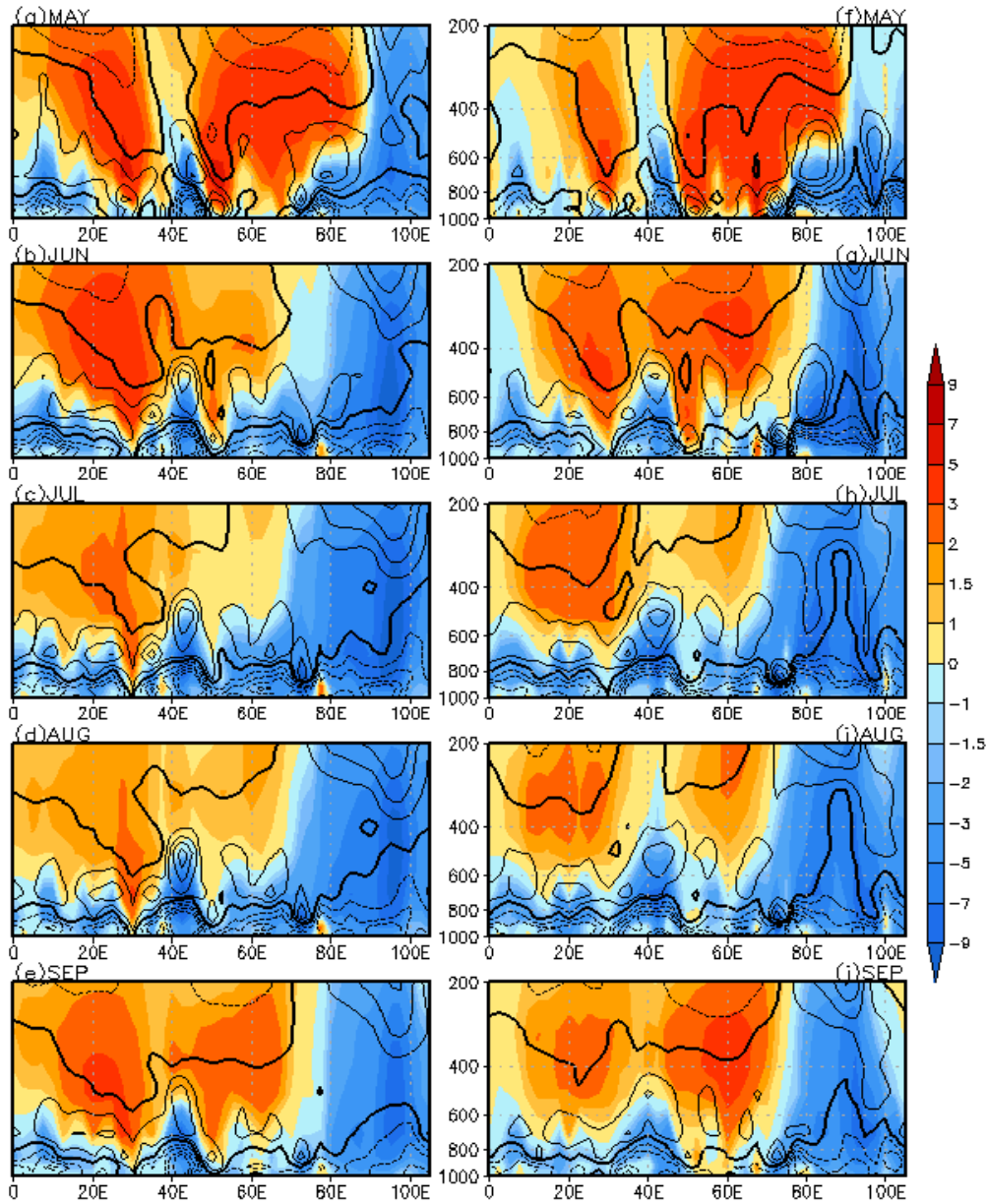


Figure 2: Seasonal evolution (May to September) of vertical component of velocity (shading, units in $10^{-2} \text{ Pa s}^{-1}$) and horizontal divergence (contours, units in 10^{-6} s^{-1}), along a pressure-longitude plane averaged over 15° - 30°N for ERAi (left panel, a-e) and CTRL (right panel, f-j). The negative (dashed) and positive (continuous) contours correspond, respectively, to absolute magnitudes of 1, 2, 3 and 4 units. The zero contours are highlighted in thick black colour.

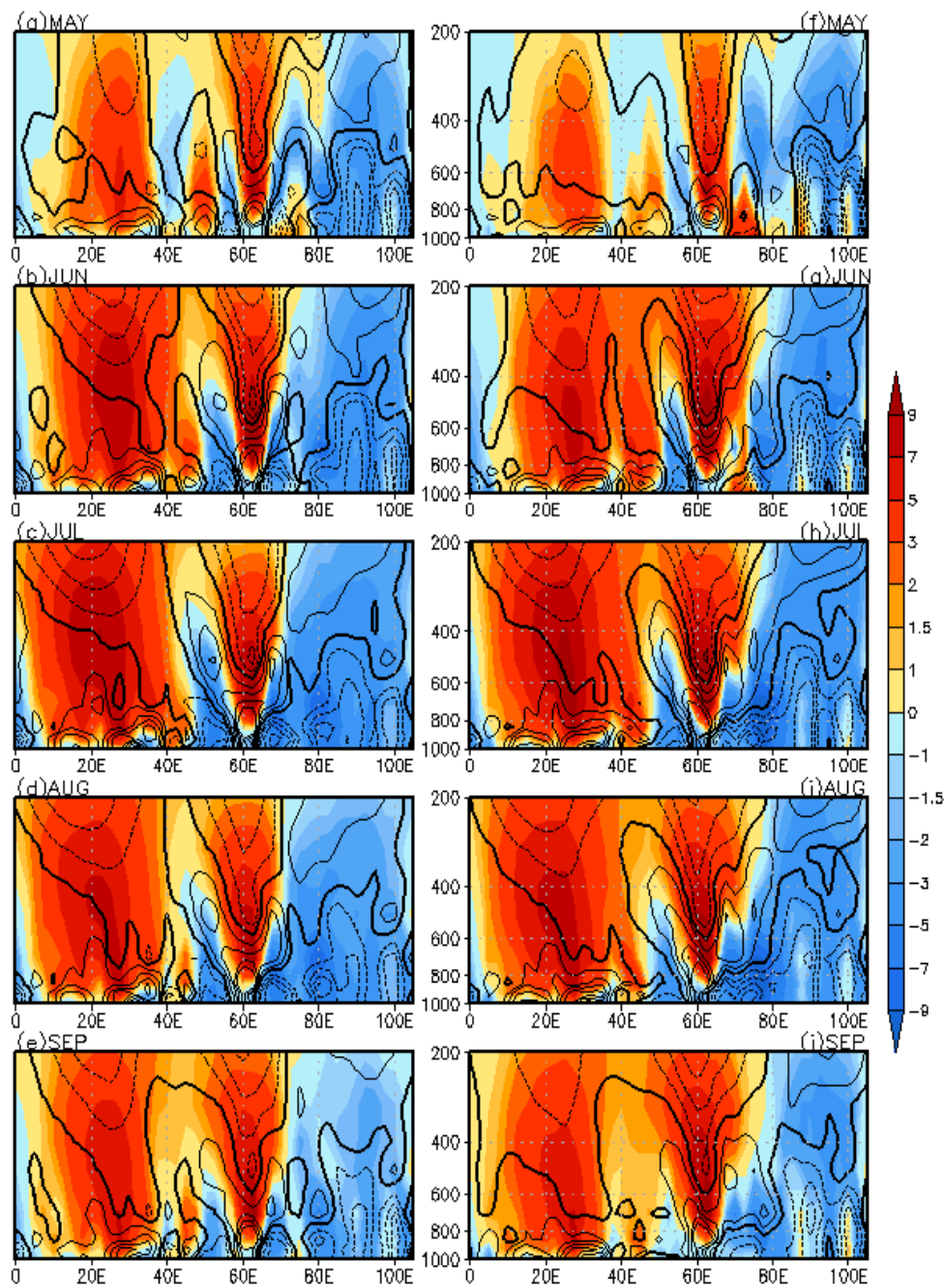


Figure 3 (a-j): Same Figure 2, but for the vertical component of velocity (shading, units in 10^{-2} Pa s^{-1}) and horizontal divergence (contours, units in 10^{-6} s^{-1}), averaged over 30° - 40° N. The negative (dashed) and positive (continuous) contours correspond, respectively, to absolute magnitudes of 1, 2, 3 and 4 units. The zero contours are highlighted in thick black colour.

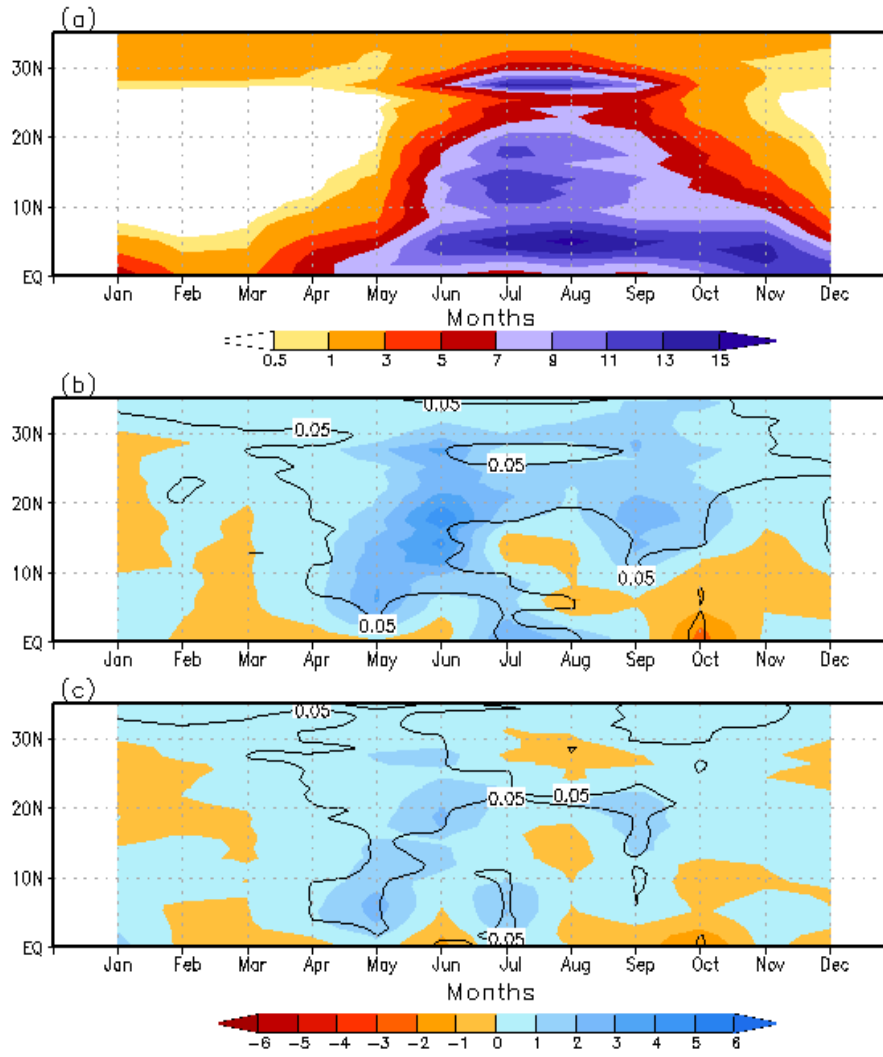


Figure 4: (a) Time-latitude evolution of rainfall climatology (mm day^{-1} , containing both land and oceanic points) averaged along 70° - 90° E, from CTRL. (b) and (c) time-latitude evolution of anomalous rainfall response (computed against the CTRL rainfall climatology, see Section 2) averaged along 70° - 90° E from Desert_m20 and from Desert_Arab_m20, respectively. In (b) and (c), the responses that are above the 95% confidence level according to a permutation procedure with 9999 shuffles are encircled.

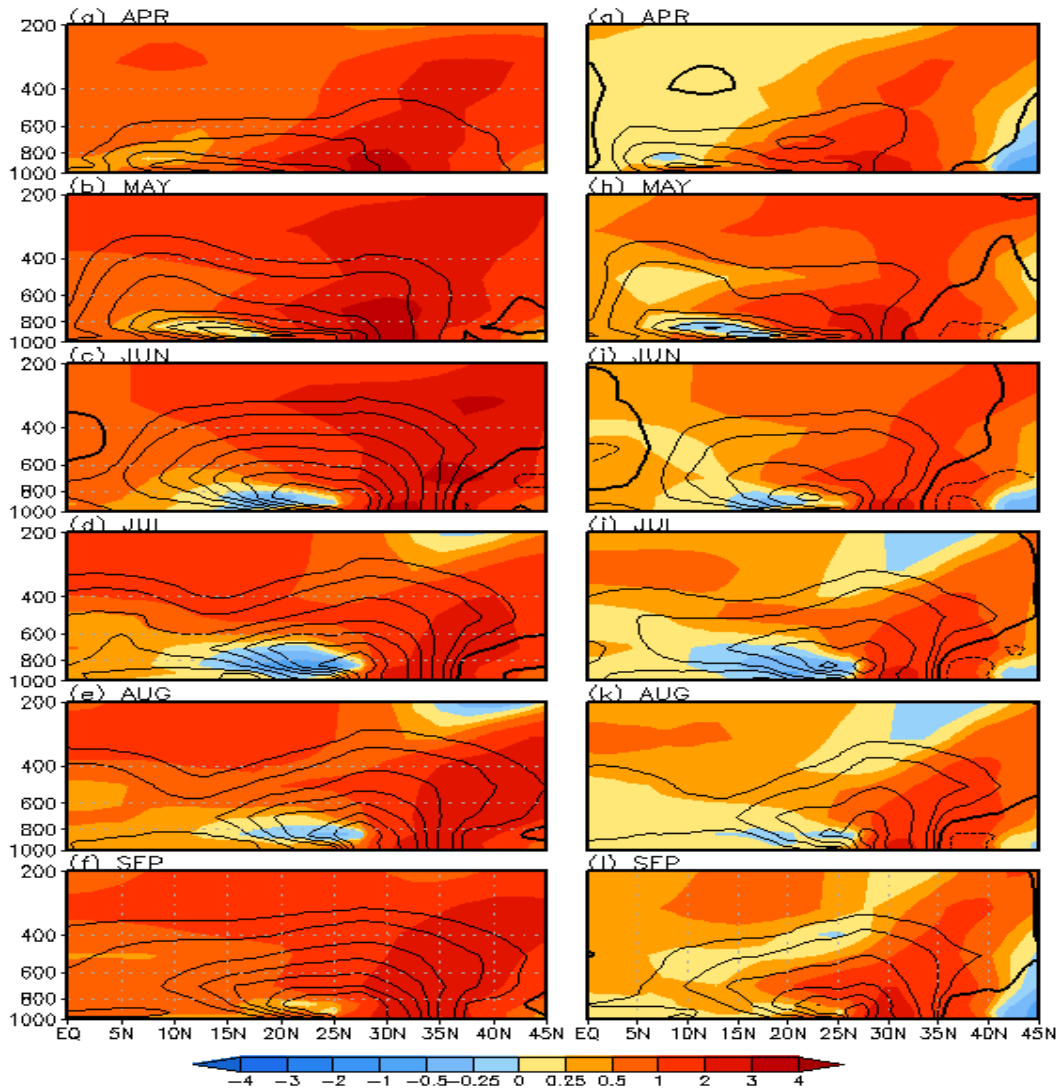


Figure 5: (Left Panel, a-f) Seasonal evolution (April to September) of response in temperature (shading, °C) and specific humidity (contours, $\times 10^{-3} \text{ Kg Kg}^{-1}$) along a pressure-latitude plane over the longitudes 55° - 75° E, for Desert_m20 experiment. Right panel (g-l), same as (a-f), but for Desert_Arab_m20 experiment. The anomalous responses are computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond, respectively, to absolute magnitudes of 0.25, 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero contours are highlighted in thick black colour.

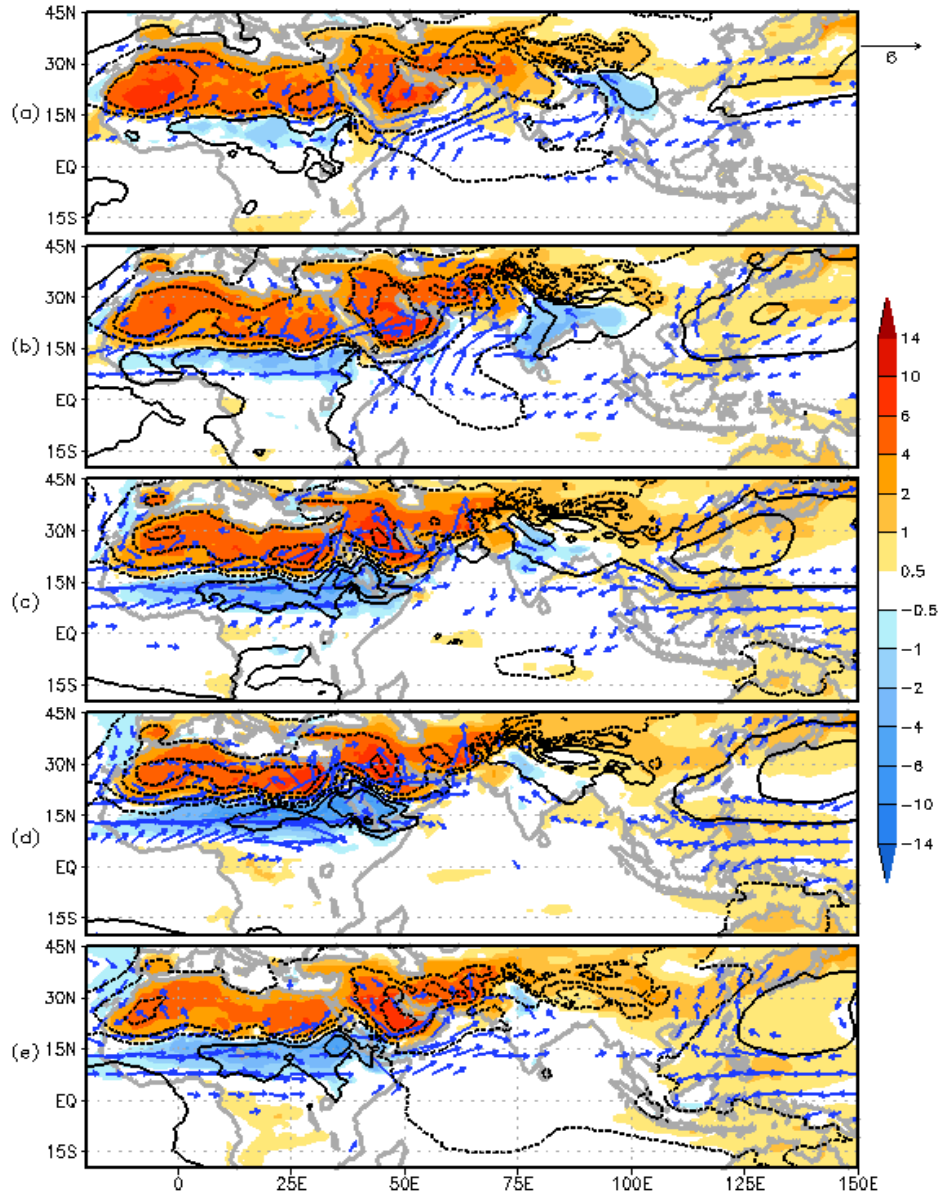


Figure 6: Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS (shading, °C) and 850-hPa wind (ms^{-1}) from Desert_m20 experiment. Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6 units. The zero contours are not drawn. The anomalous response is computed against the CTRL climatology (see Section 2). Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn.

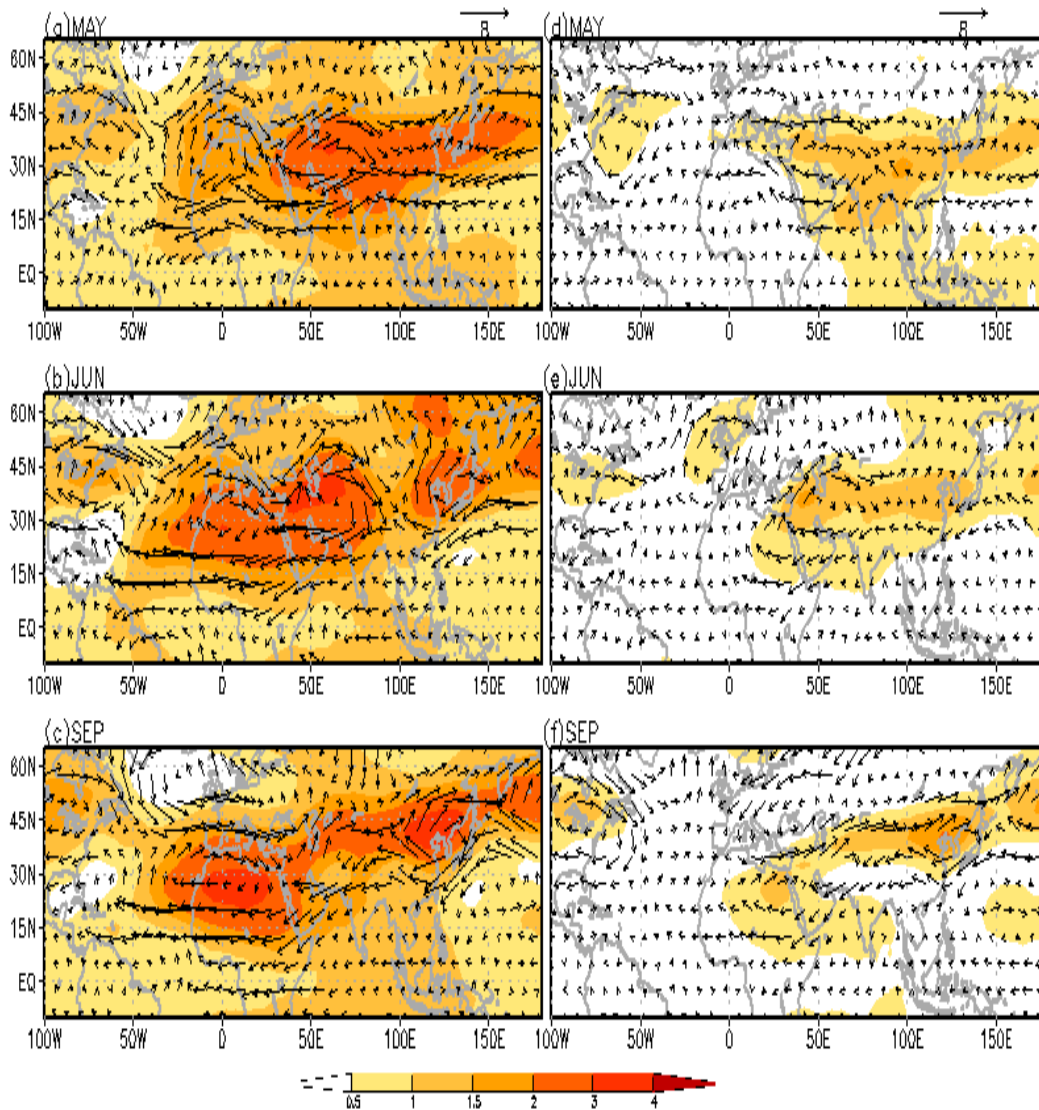


Figure 7: Map of the anomalous responses in wind and temperature at 300-hPa for (a) May, (b) June and (c) September, from Desert_m20 experiments. (d)-(f), same as (a)-(c), but from Desert_Arab_m20. The anomalous response is against the CTRL climatology (see Section 2).

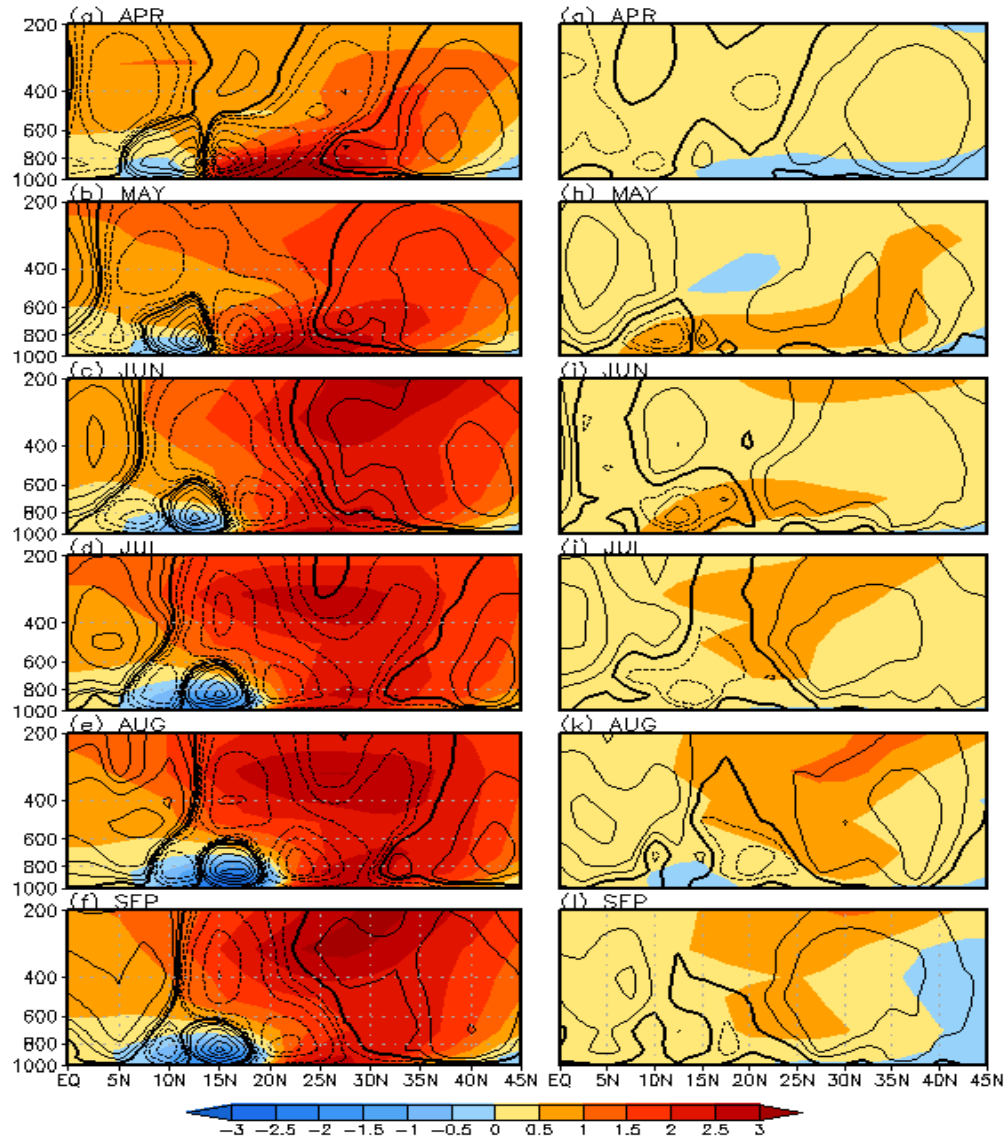


Figure 8: (Left Panel, a-f) Seasonal evolution (April to September) of response in temperature (shading, °C) and vertical component of velocity (contours, units is $10^{-2} \text{ Pa s}^{-1}$) along a pressure-latitude plane averaged over the longitudes 20°W - 30°E , for Desert_m20 experiment. Right panel (g-l), same as (a-f), but for Desert_Arab_m20 experiment. The anomalous response is computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.25, 0.5, 1, 2, 3, 4 and 5 units. The zero contours are highlighted in thick black colour.

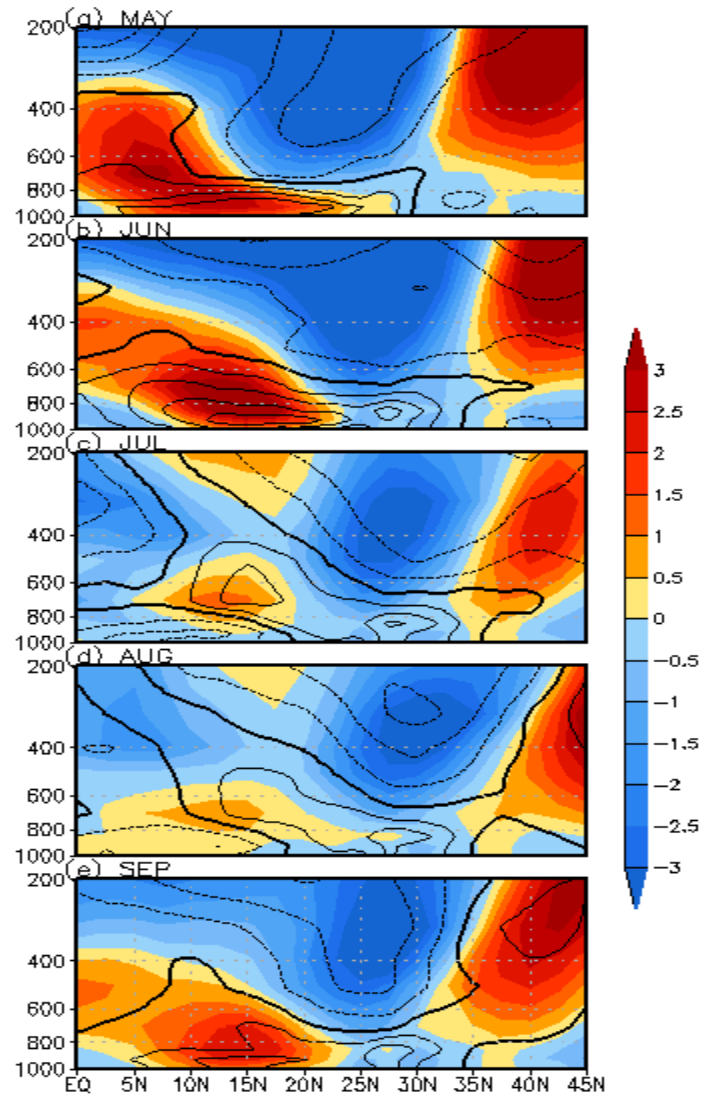


Figure 9: Seasonal evolution (May to September) of response in horizontal component of wind (zonal wind in shading and meridional wind in contours, ms^{-1}) along a pressure-latitude plane averaged over the longitudes 55° - 75°E , for Desert_m20 experiment. The anomalous response is computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero contours are highlighted in thick black colour.

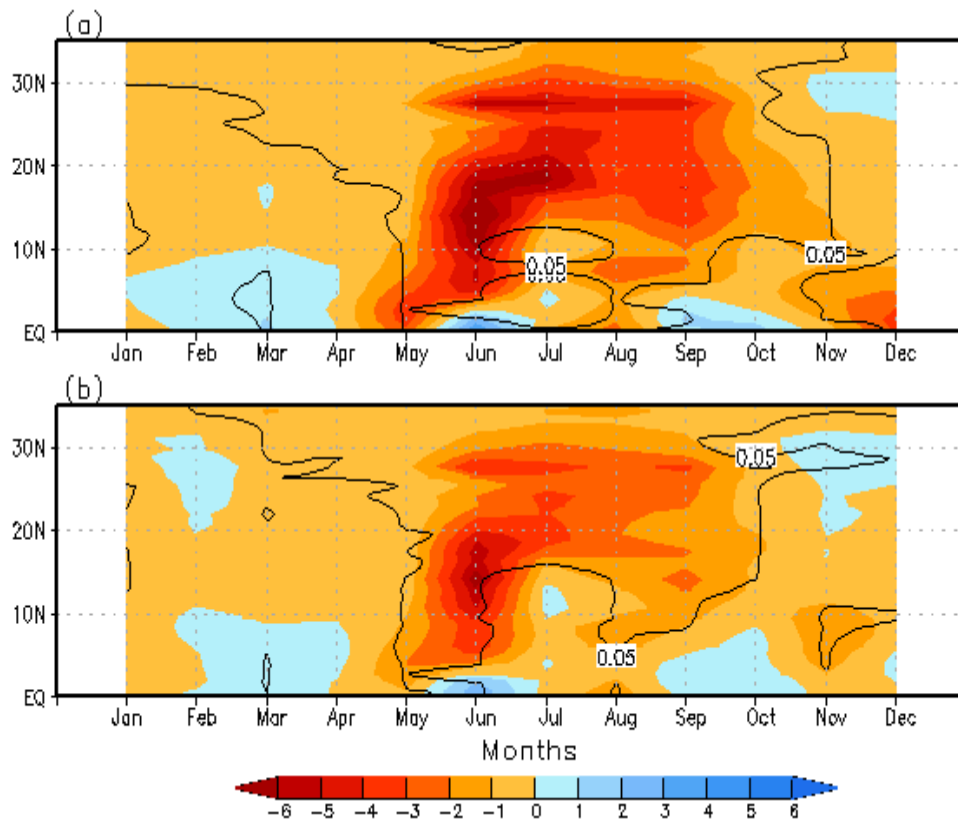


Figure 10: Time-latitude evolution of rainfall response (mm day^{-1} , averaged along $70^{\circ}\text{-}90^{\circ}\text{E}$) from the sensitivity experiments. In (a) from Desert_p20 and in (b) from Desert_Arab_p20. The anomalous responses are computed against the CTRL climatology (see Section 2). In all panels, the responses that are above the 95% confidence level according to a permutation procedure with 9999 shuffles are encircled.

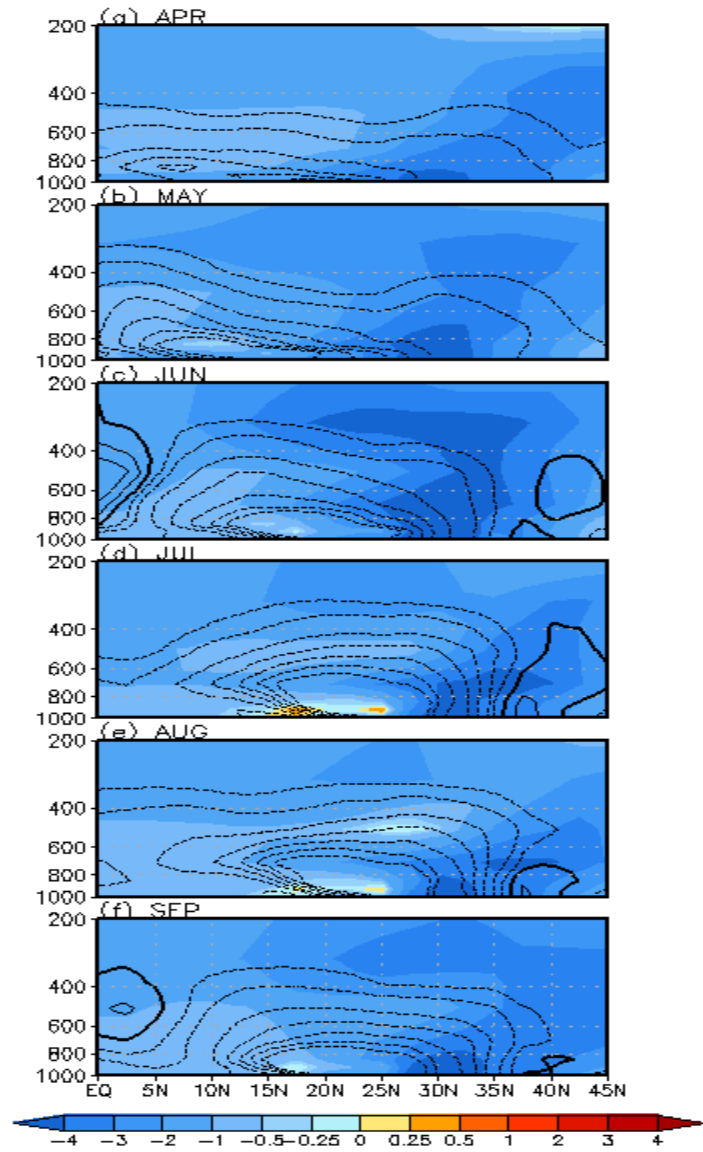


Figure 11: Seasonal evolution (April to September) of response in temperature (shading, °C) and specific humidity (contours, $\times 10^{-3} \text{ Kg Kg}^{-1}$) along a pressure-latitude plane over the longitudes 55°-75°E, for Desert_p20 experiment. The anomalous response is computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond, respectively, to magnitudes of 0.25, 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero contours are highlighted in thick black colour.

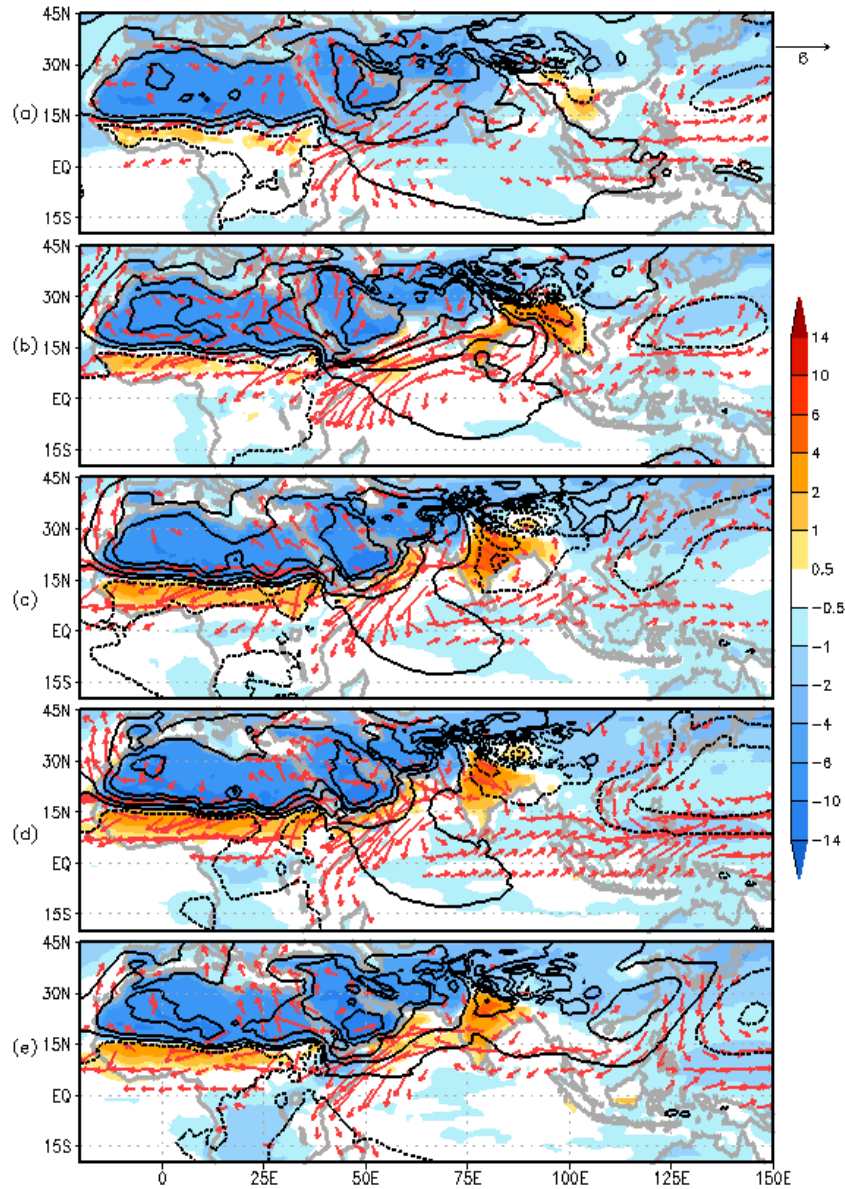


Figure 12: Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS (shading, °C) and 850-hPa wind (ms^{-1}) from Desert_p20 experiment. Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6 units. The zero contours are not drawn. The anomalous response is computed against the CTRL climatology (see Section 2). Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn.

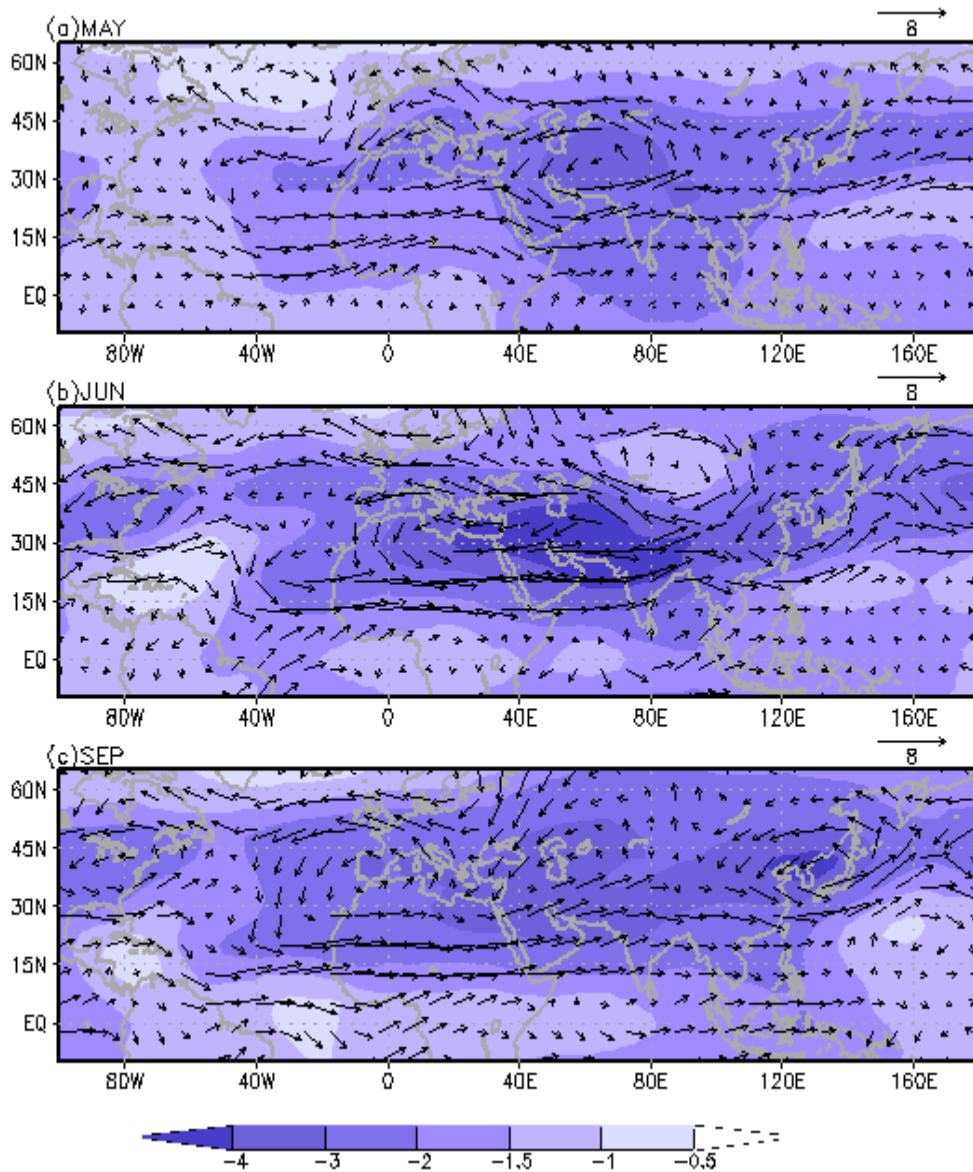


Figure 13: (a) Anomalous response in wind (ms^{-1}) and temperature ($^{\circ}\text{C}$, shading) at 300-hPa, during May from Desert_p20 experiment. (b) and (c), same as (a), but for June and September, respectively. The anomalous response is computed against the CTRL climatology (see Section 2).

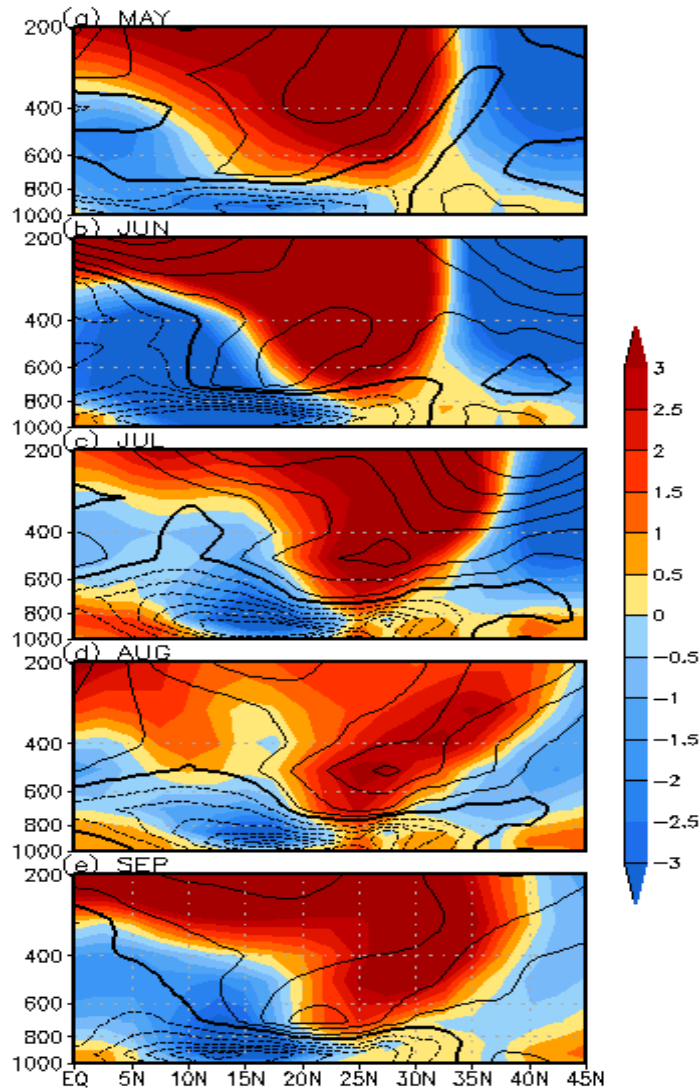


Figure 14: Seasonal evolution (May to September) of response in horizontal component of wind (zonal wind in shading and meridional wind in contours, ms^{-1}) along a pressure-latitude plane averaged over the longitudes $55^{\circ}\text{-}75^{\circ}\text{E}$, for Desert_p20 experiment. The anomalous response is computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond, respectively, to magnitudes of 0.5, 1, 1.5, 2, 2.5 and 3 units. The zero contours are highlighted in thick black colour.

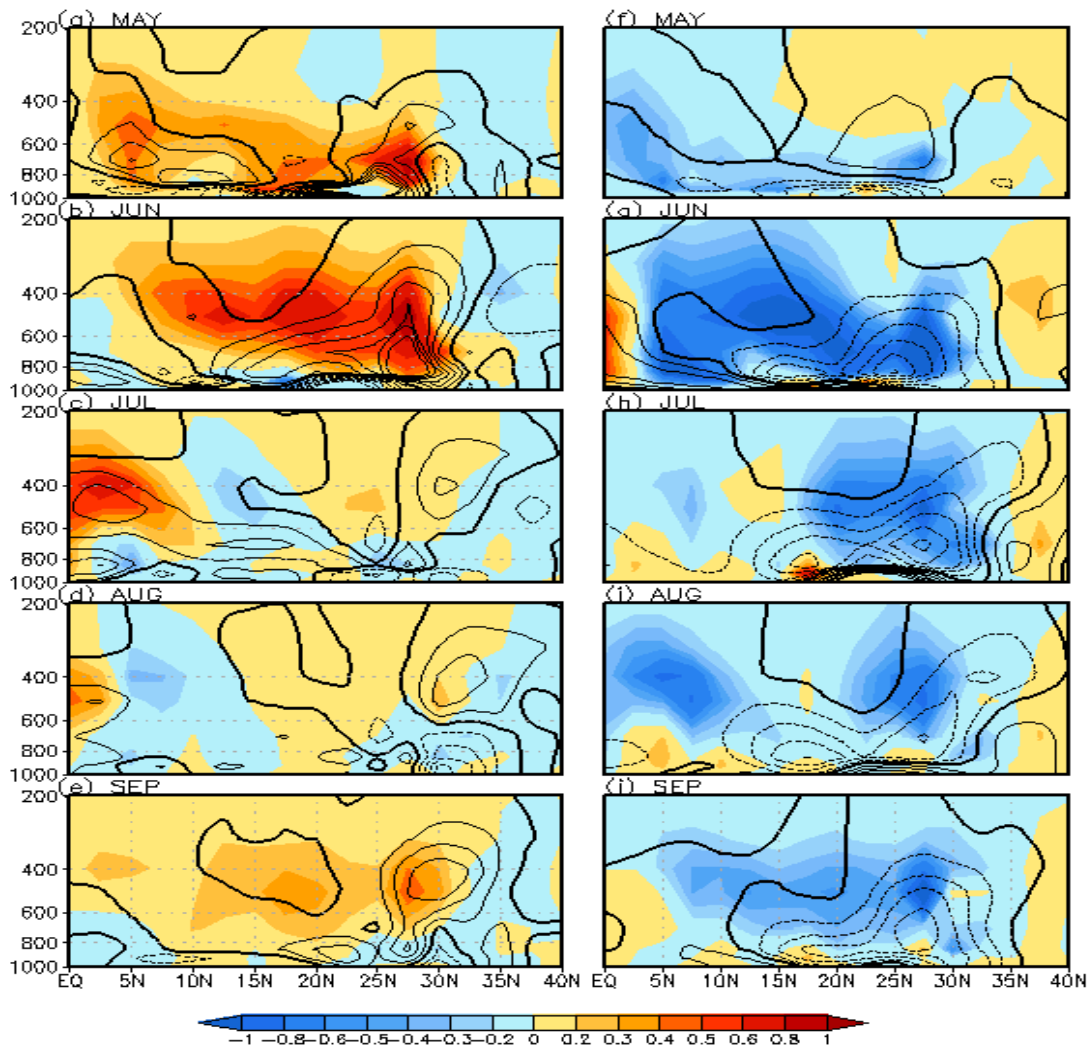


Figure 15: (Left Panel, a-e) Seasonal evolution (May to September) of response in vertical moisture advection (shading, units is $10^{-3} \text{ Kg Kg}^{-1} \text{ day}^{-1}$) and horizontal moisture advection (contours, $\times 10^{-3} \text{ Kg Kg}^{-1} \text{ day}^{-1}$) along a pressure-latitude plane over the longitudes 75° - 90° E, for Desert_m20 experiment. Right panel (f-j), same as (a-e), but for Desert_p20 experiment. The anomalous response is computed against the CTRL climatology (see Section 2). The negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.1, 0.2, 0.3, 0.4, 0.5 and 0.6 units. The zero contours are highlighted in thick black colour.

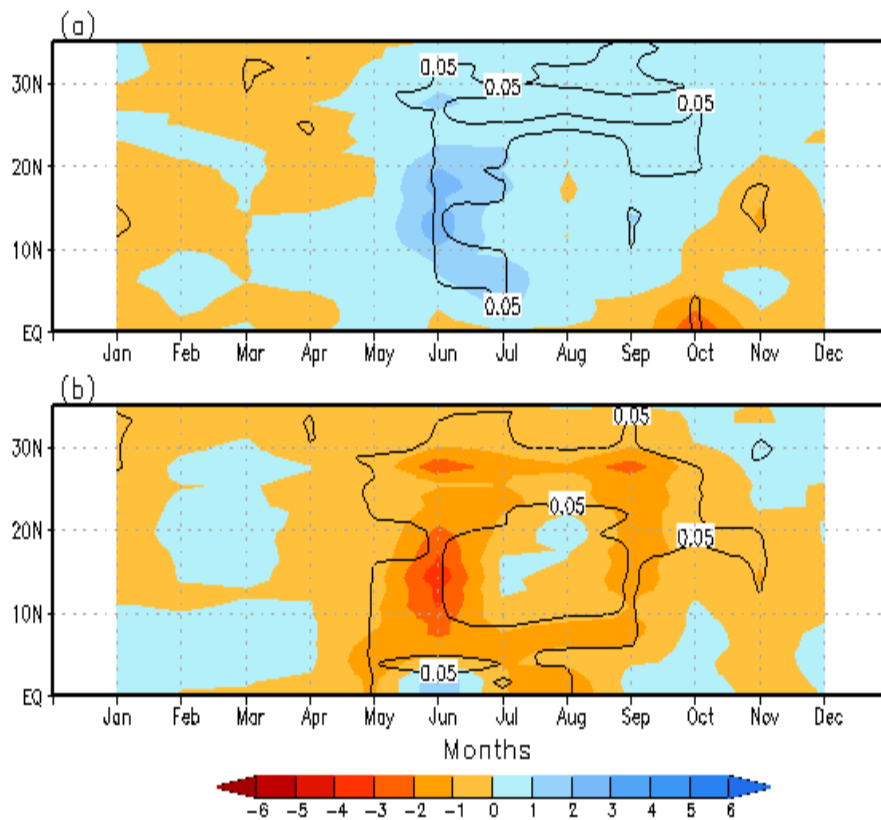


Figure 16: (a) Time-latitude evolution of rainfall response (mm day^{-1} , averaged along 70° - 90°E) from the Desert_Sahara_m20 experiment. In (b), same as (a) but for Desert_Sahara_p20 experiment. The anomalous responses are computed against the CTRL climatology (see Section 2). In (a) and (b), the responses that are above the 95% confidence level according to a permutation procedure with 9999 shuffles are encircled.

Supplementary Figures:

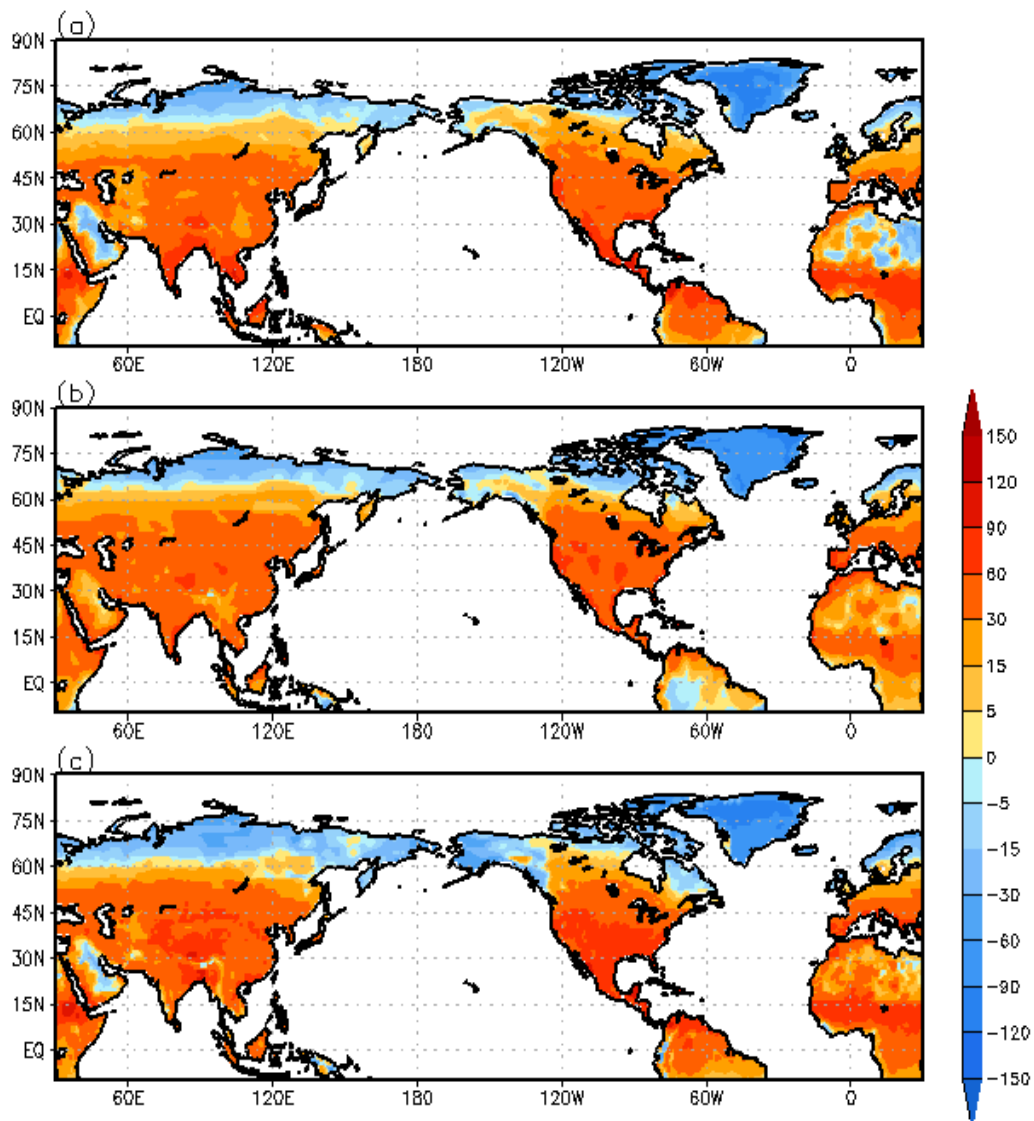


Figure S1: Climatological map of net radiation budget at TOA (Wm^{-2} , see Table 2 and Section 2 for more details on radiation budget) for June to September period. In (a). CERES-EBAF, (b) ERAi and (c) CTRL

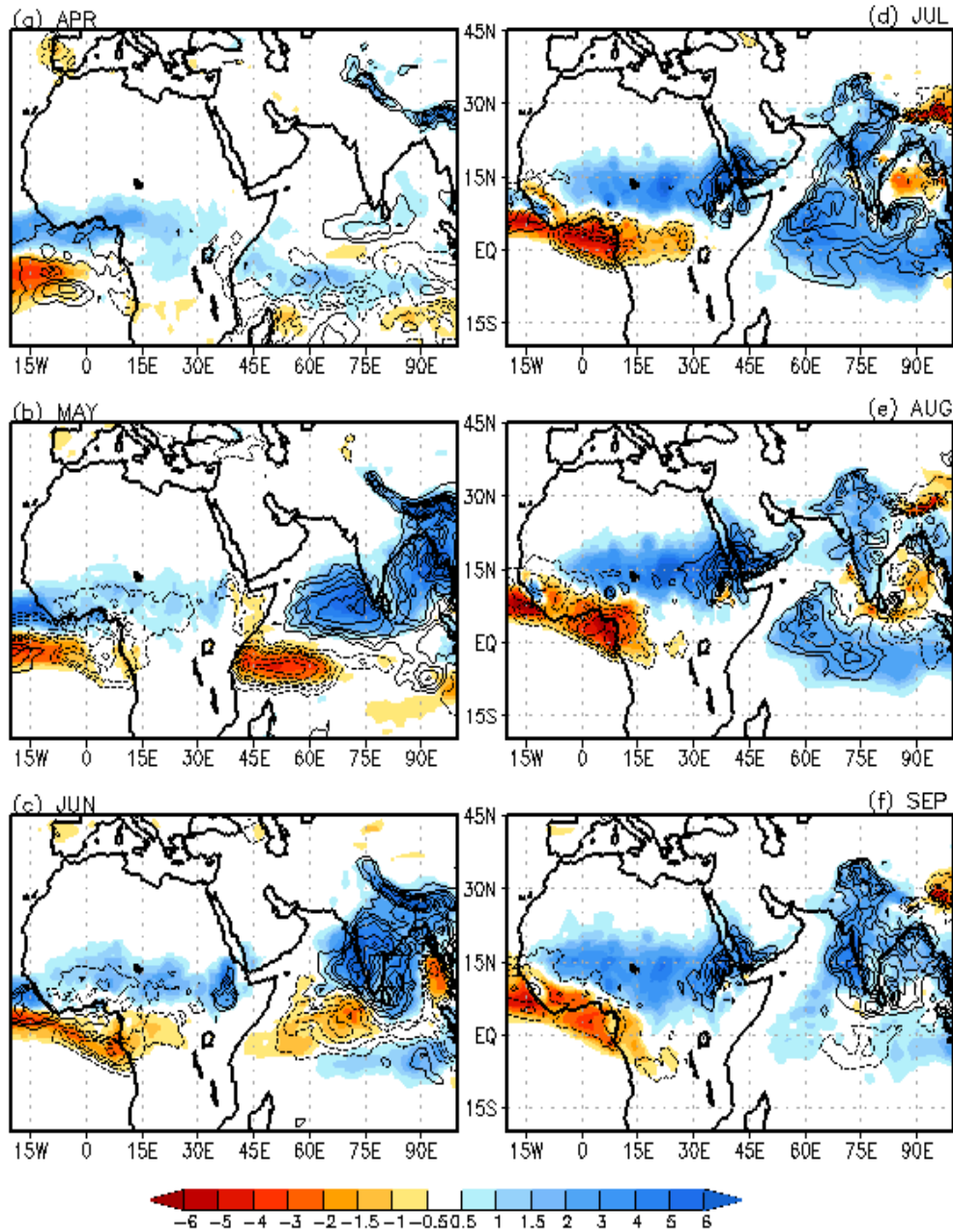


Figure S2: Spatio-temporal evolution of rainfall (May to September, mm day⁻¹) response from the Desert_m20 experiment (in shading). The overlying contours are the rainfall response from Desert_Arab_m20. Negative (dashed) and positive (continuous) contours are drawn following the shading convention. The anomalous responses are computed against the CTRL climatology (see Section 2).

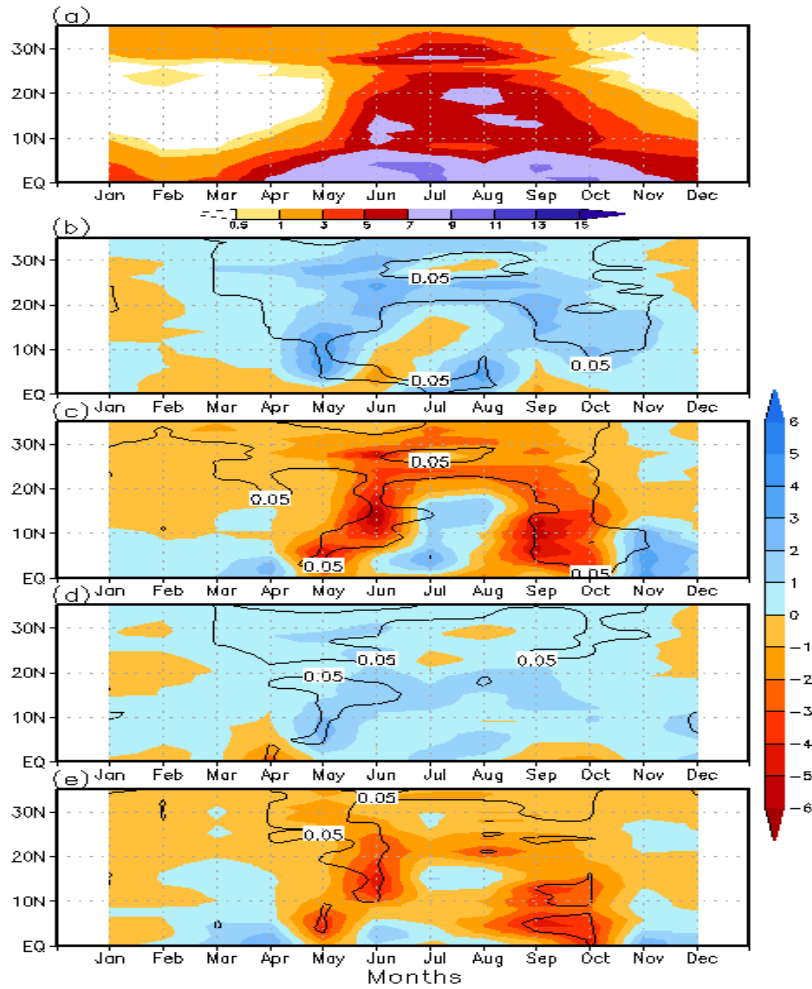


Figure S3: In (a) to (c), same as that of Figure 4, but for the various experiments (CTRL, Desert_m20 and Desert_p20), conducted using CFSv2 coupled model (see Terray et al. 2017). Similarly, in (d) and (e), same as that of Figure 11, but for Desert_Arab_m20 and Desert_Arab_p20 simulations, using CFSv2. The length of integration (i.e. using CFSv2) in CTRL simulation is for 40 years, while it is 20 years for the Desert_m20 and Desert_p20. Note that the configuration used for the experiments also takes advantage of up-to-date satellite MODIS data for estimating the background snow-free albedo as described in Terray et al. (2017). The first 10 years of all the simulations are excluded for all the analyses presented in the text. In (b) and (c), the responses that are above the 95% confidence level according to a permutation procedure with 9999 shuffles are encircled.

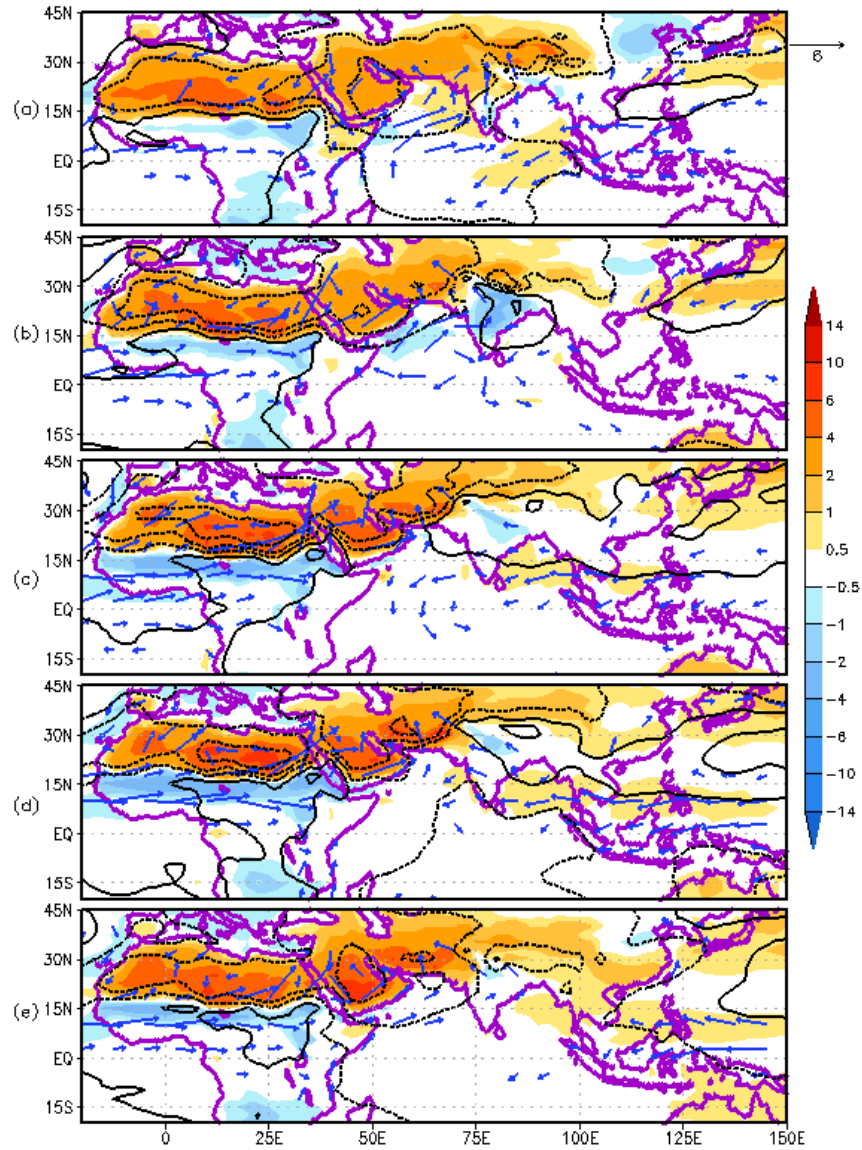


Figure S4: Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS (shading, °C) and 850-hPa wind (ms^{-1}) from Desert_m20 experiment done using CFSv2 model. Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6 units. The anomalous response is computed against the CTRL climatology. Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn. For more details of the CFS runs, please refer to the legend of Fig. S3.

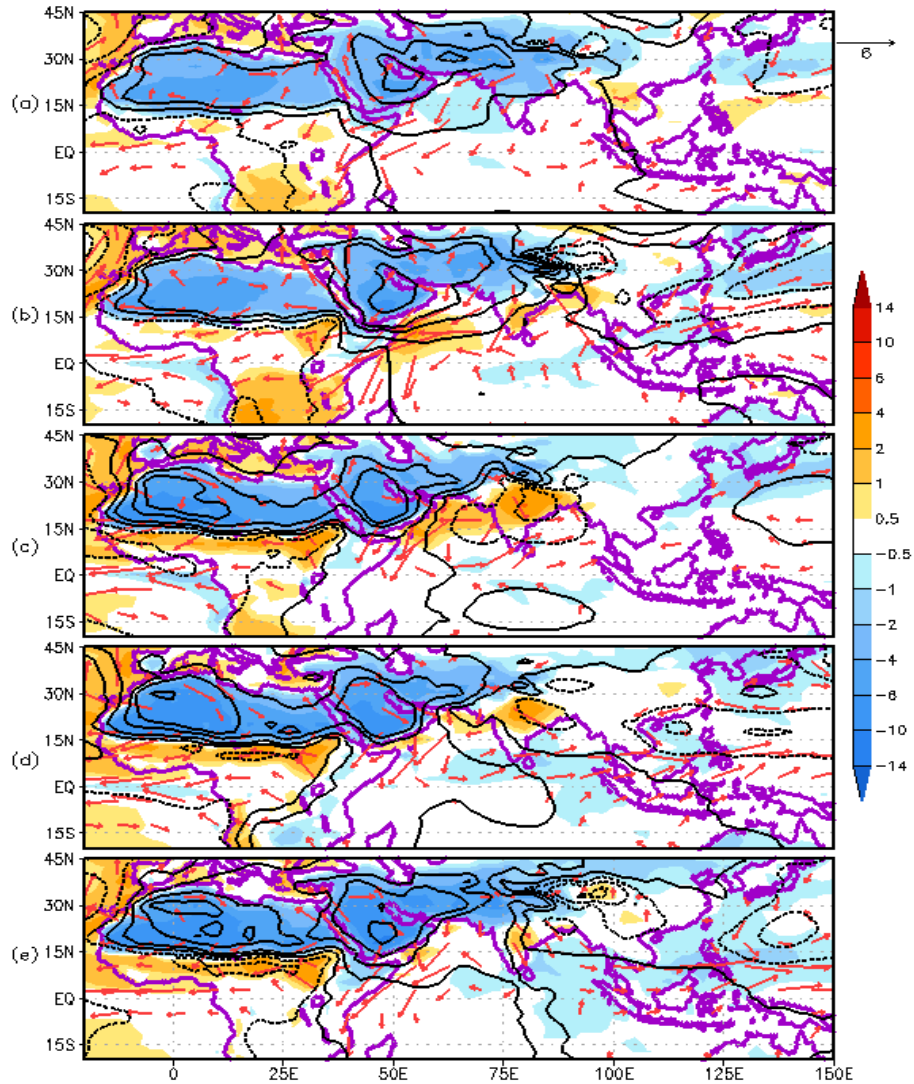


Figure S5: Seasonal evolution (May to September) of response in MSLP (contours, hPa), TS (shading, °C) and 850-hPa wind (ms^{-1}) from Desert_p20 experiment done using CFSv2 model. Negative (dashed) and positive (continuous) contours correspond respectively to magnitudes of 0.5, 1.5, 3, 4 and 6 units. The anomalous response is computed against the CTRL climatology. Wind vectors of magnitudes exceeding 1 ms^{-1} are only drawn. For more details of the CFS runs, please refer to the legend of Fig. S3.

Table 1: Details of coupled model experiments as performed using SINTEX-F2 coupled model. The first 10 years of all the simulations are excluded for all the analyses presented in the text. See Section 2 for more details of the experiments.

Experiment name	Length of integration (years)	Experimental setup
CTRL	110	Control experiment using background snow-free broadband shortwave albedo prescribed from MODIS products (Terray et al. 2017)
Desert_m20	60	Similar to CTRL with the exception that the background land MODIS albedo has been artificially decreased by -20% over the hot subtropical desert (15°-40°N, 20°W-75°E)
Desert_p20	60	Same as Desert_m20, but with an increase in background albedo of +20% over the hot subtropical desert (15°-40°N, 20°W-75°E)
Desert_Arab_m20	60	Similar to CTRL, but with a decrease in background albedo of -20% over the Arabia and Middle-East deserts only (15°-40°N, 35°E-75°E)
Desert_Arab_p20	60	Same as Desert_Arab_m20, but with an increase in background albedo of +20% over the Arabia and Middle-East deserts (15°-40°N, 35°E-75°E)
Desert_Sahara_m20	60	Similar to CTRL, but with a decrease in background albedo of -20% over the Sahara desert only (15°-40°N, 20°W-35°E)
Desert_Sahara_p20	60	Same as Desert_Sahara_m20, but with an increase in background albedo of +20% over the Sahara desert (15°-40°N, 20°W-35°E)

Table 2: The radiative and turbulent fluxes, at surface (SURF) and top of the atmosphere (TOA), averaged over the land points in the domain 15°-40°N and 20°W-75°E (e.g. hot subtropical desert region) for June to September (JJAS) season. Here, SW and LW are shortwave and long-wave radiations, respectively. Net Rad means net radiation following the formulation in Su and Neelin (2002). SH and LH are sensible and latent heat fluxes. TS and PR stands for surface skin temperature and rainfall, respectively. Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). Asterisk (*) in the Table represents that the rainfall (PR) is averaged using the GPCP data and not using ERAi estimates. All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and $mm\ day^{-1}$, respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LW at TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR*
ERAi	26 (28)	330 (211)	-300 (-494)	-300 (-119)	30 (92)	-19	-69	305.8	0.5
CERES- EBAF	29 (29)	310 (206)	-294 (-501)	-294 (-111)	16 (95)				
CTRL	28 (30)	323 (202)	-294 (-497)	-294 (-116)	29 (86)	-9	-73	305.6	0.3

Table 3: The radiative and turbulent fluxes averaged over Indian landmass (land points only: 75°-90°E, 5°-25°N) for JJAS season. Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). Asterisk (*) in Table represents that the rainfall (PR) is averaged using the GPCP data and not using ERAi estimates. All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and $mm\ day^{-1}$, respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LW at TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR*
ERAi	34 (17)	290 (155)	-237 (-460)	-237 (-39)	53 (116)	-92	-27	300.4	7.7
CERES- EBAF	35 (14)	284 (167)	-216 (-461)	-216 (-36)	68 (131)				
CTRL	39 (16)	269 (154)	-223 (-474)	-223 (-41)	46 (113)	-68	-47	302.3	6

Table 4: The radiative and turbulent fluxes averaged over the land points in the hot subtropical desert domain (15°-40°N and 20°W-75°E) during JJAS season for the negative albedo perturbation experiments. Asterisk (*) in Table indicates that the flux values are averaged only over 15°-40°N and 35°-75°E (e.g. in the Desert_Arab_m20 experiment). Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). The upper part of the Table provides the climatological estimates from the sensitivity experiments while the lower part shows the anomalous responses (deviations from the CTRL climatology, see Section 2). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and $mm\ day^{-1}$, respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LW at TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR
	Climatological responses								
Desert_m20	22 (11)	348 (233)	-280 (-516)	-280 (-110)	68 (123)	-17	-102	308.5	0.8
*Desert_Arab_m20	19 (7)	361 (255)	-293 (-509)	-293 (-119)	68 (136)	-16	-115	307.3	0.6
	Anomalous responses								
Desert_m20	-6 (-19)	25 (31)	14 (-19)	14 (6)	39 (37)	-8	-29	2.9	0.5
*Desert_Arab_m20	-7 (-19)	30 (36)	10 (-17)	10 (0)	40 (36)	-2	-33	2.6	0.25

Table 5: The radiative and turbulent fluxes averaged over Indian landmass (land points only: 75°-90°E, 5°-25°N) during JJAS season for the negative albedo perturbation experiments. Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). The upper part of the Table provides the climatological estimates from the sensitivity experiments while the lower part shows the anomalous responses (deviations from the CTRL climatology). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and mm day^{-1} , respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LWat TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR
	Climatological responses								
Desert_m20	42 (16)	256 (138)	-214 (-470)	-214 (-32)	42 (106)	-78	-31	301.6	8
Desert_Arab _m20	41 (16)	261 (145)	-219 (-471)	-219 (-36)	42 (109)	-72	-37	302	7
	Anomalous responses								
Desert_m20	3 (0)	-13 (-16)	9 (4)	9 (9)	-4 (-7)	-10	16	-0.7	2
Desert_Arab _m20	2 (0)	-8 (-9)	4 (3)	4 (5)	-4 (-4)	-4	10	-0.3	1

Table 6: The radiative and turbulent fluxes averaged over the hot subtropical desert domain during JJAS season for the positive albedo perturbation experiments. Asterisk (*) in Table indicates that the flux values are averaged only over 15°-40°N and 35°-75°E (e.g. in the Desert_Arab_p20 experiment). Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). The upper part of the Table provides the climatological estimates from the sensitivity experiments while the lower part shows the anomalous responses (deviations from the CTRL climatology). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and $mm\ day^{-1}$, respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LW at TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR
	Climatological responses								
Desert_p20	37 (51)	283 (156)	-293 (-454)	-293 (-110)	-10 (46)	-7	-35	298.8	0.1
*Desert_Arab_p20	35 (46)	289 (171)	-299 (-460)	-299 (-111)	-9 (60)	-11	-44	299.7	0.2
	Anomalous responses								
Desert_p20	9 (21)	-40 (-46)	1 (43)	1 (6)	-39 (-40)	2	38	-6.8	-0.2
*Desert_Arab_p20	9 (2)	-42 (-48)	4 (32)	4 (8)	-37 (-40)	3	38	-5	-0.15

Table 7: The radiative and turbulent fluxes averaged over Indian landmass (land points only: 75°-90°E, 5°-25°N) during JJAS season for the positive albedo perturbation experiments. Fluxes are signed in the direction of the fluxes (i.e. negative sign indicates upward fluxes and vice-versa). The upper part of the Table provides the climatological estimates from the sensitivity experiments while the lower part shows the anomalous responses (deviations from the CTRL climatology, see Section 2). All fluxes have energy units (Wm^{-2}) and albedo is in %, while TS and PR have the units of K and $mm\ day^{-1}$, respectively.

	Albedo at TOA (SURF)	Net SW at TOA (SURF)	LWat TOA (SURF)	Net LW at TOA (SURF)	Net Rad at TOA (SURF)	LH Flux	SH flux	TS	PR
	Climatological responses								
Desert_p20	31 (16)	306 (199)	-251 (-487)	-251 (-70)	55 (129)	-44	-85	304.3	2.8
Desert_Arab_p20	34 (16)	294 (184)	-241 (-483)	-241 (-59)	53 (125)	-53	-72	303.6	3.8
	Anomalous responses								
Desert_p20	-8 (0)	37 (45)	-28 (-13)	-28 (-29)	9 (16)	24	-38	2	-3.2
Desert_Arab_p20	-5 (0)	25 (30)	-18 (-9)	-18 (-18)	7 (12)	15	-25	1.3	-2.2